

Expectations from T2K & NOvA in light of recent Neutrino Data

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References

Based on the papers:

- 1) Potential of optimized NOvA for large θ_{13} & combined performance with a LArTPC and T2K*

SKA, S. Prakash, S. Raut, S. Uma Sankar, arXiv: 1208.3644

- 2) Resolving the Octant of θ_{23} with T2K and NOvA*

SKA, S. Prakash, S. Uma Sankar, arXiv: 1301.2574

Global Analysis of World Neutrino Data

Reference	Ref. [53]	Ref. [54]	Ref. [55]
$\Delta m_{21}^2 (10^{-5} \text{eV}^2)$	7.62 ± 0.19	$7.54^{+0.26}_{-0.22}$	$7.50^{+0.205}_{-0.160}$
$\Delta m_{31}^2 (10^{-3} \text{eV}^2)$ (N)	$2.55^{+0.06}_{-0.09}$	$2.43^{+0.06}_{-0.10}$	$2.49^{+0.055}_{-0.051}$
$\Delta m_{13}^2 (10^{-3} \text{eV}^2)$ (I)	$2.43^{+0.07}_{-0.06}$	$2.42^{+0.07}_{-0.11}$	$\Delta m_{23}^2 (10^{-3} \text{eV}^2) = 2.47^{+0.064}_{-0.073}$
$\sin^2 \theta_{13}$ (N)	$0.0246^{+0.0029}_{-0.0028}$	0.0241 ± 0.0025	0.025 ± 0.0023
$\sin^2 \theta_{13}$ (I)	$0.0250^{+0.0026}_{-0.0027}$	$0.0244^{+0.0023}_{-0.0025}$	
$\sin^2 \theta_{23}$ (N)	$0.427^{+0.034}_{-0.027} \oplus 0.613^{+0.022}_{-0.040}$	$0.386^{+0.024}_{-0.021}$	$0.41^{+0.030}_{-0.029} \oplus 0.60^{+0.020}_{-0.026}$
3σ range	$0.36 \rightarrow 0.68$	$0.331 \rightarrow 0.637$	$0.34 \rightarrow 0.67$
$\sin^2 \theta_{23}$ (I)	$0.600^{+0.026}_{-0.031}$	$0.392^{+0.039}_{-0.022}$	
3σ range	$0.37 \rightarrow 0.67$	$0.335 \rightarrow 0.663$	
$\sin^2 \theta_{12}$	$0.320^{+0.016}_{-0.017}$	$0.307^{+0.018}_{-0.016}$	0.31 ± 0.013

SkA, Francesco Lombardi, T. Takeuchi, arXiv1207.3492

Ref. 53: Forero, Tortola, Valle, arXiv:1205.4018

Ref. 54: Fogli, Lisi, Marrone, Montanino, Palazzo, Rotunno, arXiv:1205.5254

Ref. 55: Gonzalez-Garcia, Maltoni, Salvado, Schwetz, arXiv:1209.3023

Event/Background Rates in NO ν A/LArTPC

Channels	Old NO ν A (15 kt)		New NO ν A (14 kt)		LArTPC (10 kt)	
App.	Signal	Background	Signal	Background	Signal	Background
	CC	(Int+Mis-ID+NC)	CC	(Int+Mis-ID+NC)	CC	(Int+Mis-ID+NC)
$P_{\mu e}$ (NH)	62	6+1+4= 11	92	8+5+19= 32	123	18+5+7= 30
$P_{\mu e}$ (IH)	36	6+1+4= 11	54	8+5+19= 32	72	19+5+7= 31
$P_{\bar{\mu} e}$ (NH)	26	6+<1+6= 12	30	5+<1+10= 15	28	17+2+2= 21
$P_{\bar{\mu} e}$ (IH)	34	5+<1+6= 11	38	5+<1+10= 15	36	14+2+2= 18
Disapp.	Signal	Background	Signal	Background	Signal	Background
	CCQE	(NC only)	CCQE	(NC+Wrong-Sign muon)	CCQE	(NC+Wrong-Sign muon)
$P_{\mu\mu}$ (NH)	173	2	134	1+6= 7	403	7+20= 27
$P_{\mu\mu}$ (IH)	173	2	134	1+6= 7	402	7+20= 27
$P_{\bar{\mu}\bar{\mu}}$ (NH)	102	1	43	<1+18= 18	136	2+54= 56
$P_{\bar{\mu}\bar{\mu}}$ (IH)	103	1	43	<1+18= 18	137	2+54= 56

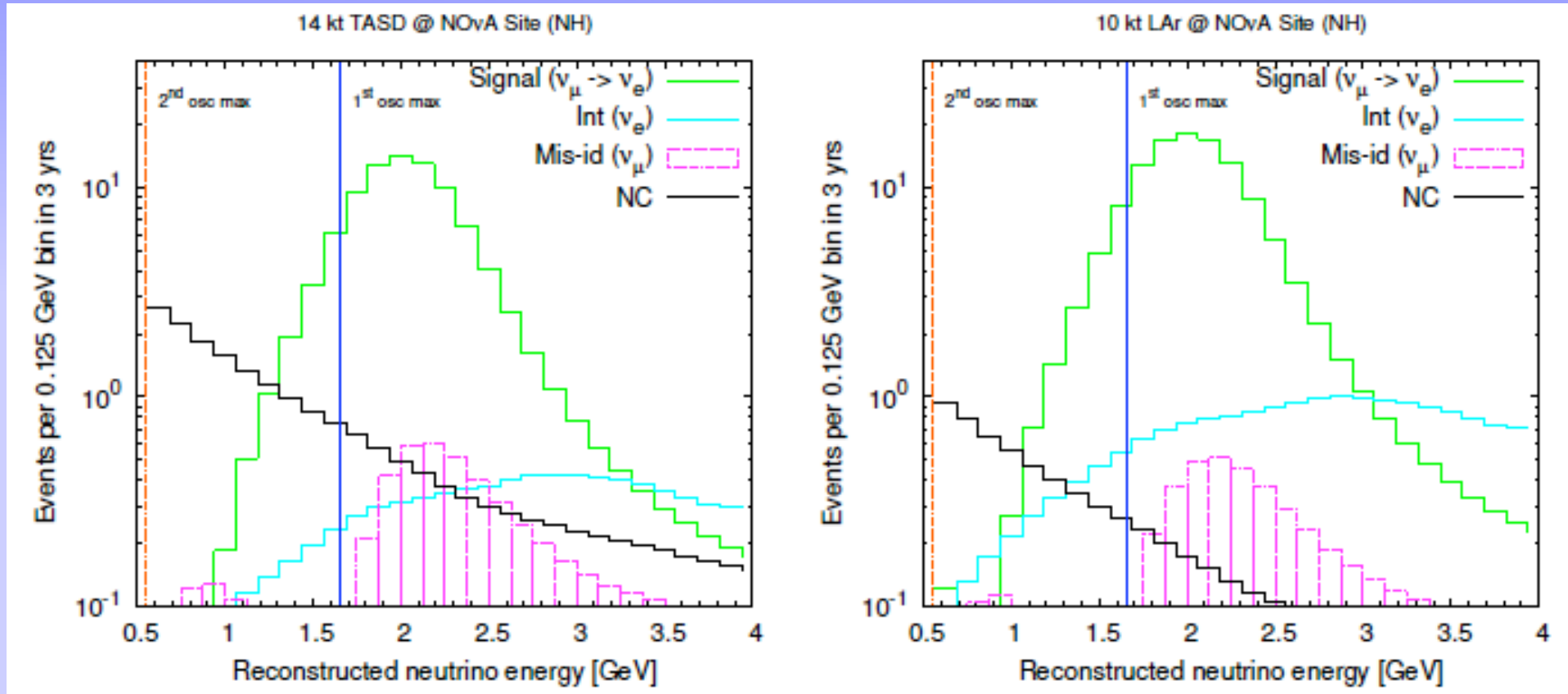
SkA, Prakash, Raut, Uma Sankar, arXiv1208.3644

NO ν A has re-optimized its event selection criteria, with more events in both signal and background

R. Patterson, Neutrino 2012 talk

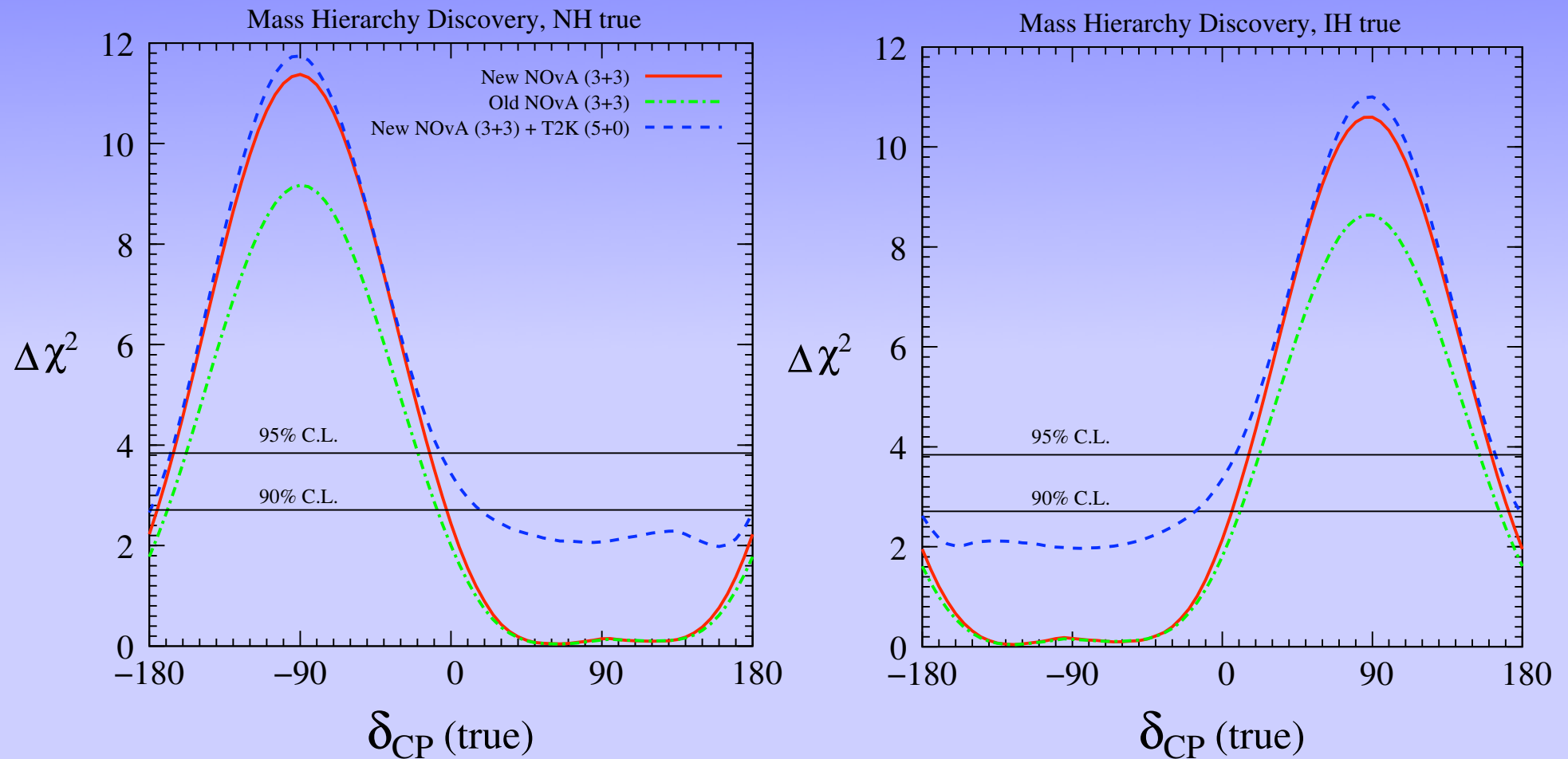
Rise in signal by 50%, background rises by a factor of 3, but we can tolerate that for large θ_{13}

Event/Background distribution in NOvA/LArTPC



SkA, Prakash, Raut, Uma Sankar, arXiv1208.3644

Present Generation Experiments: T2K and NOvA

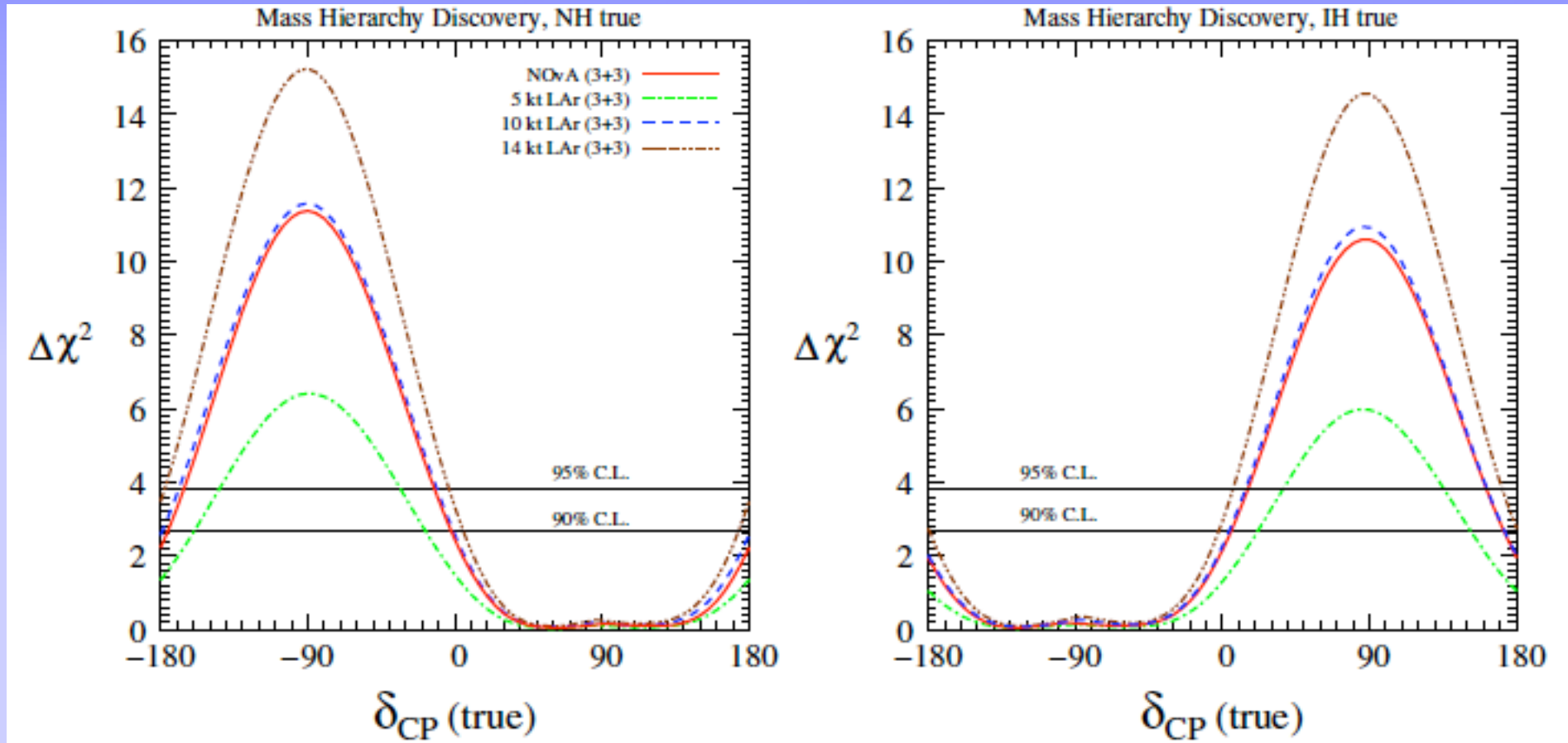


Agarwalla, Prakash, Raut, Uma Sankar, arXiv:1208.3644

Adding data from T2K and NOvA is useful to kill the intrinsic degeneracies

55% CP coverage @ 90% C.L. and 45% CP coverage @ 95% C.L. for MH discovery

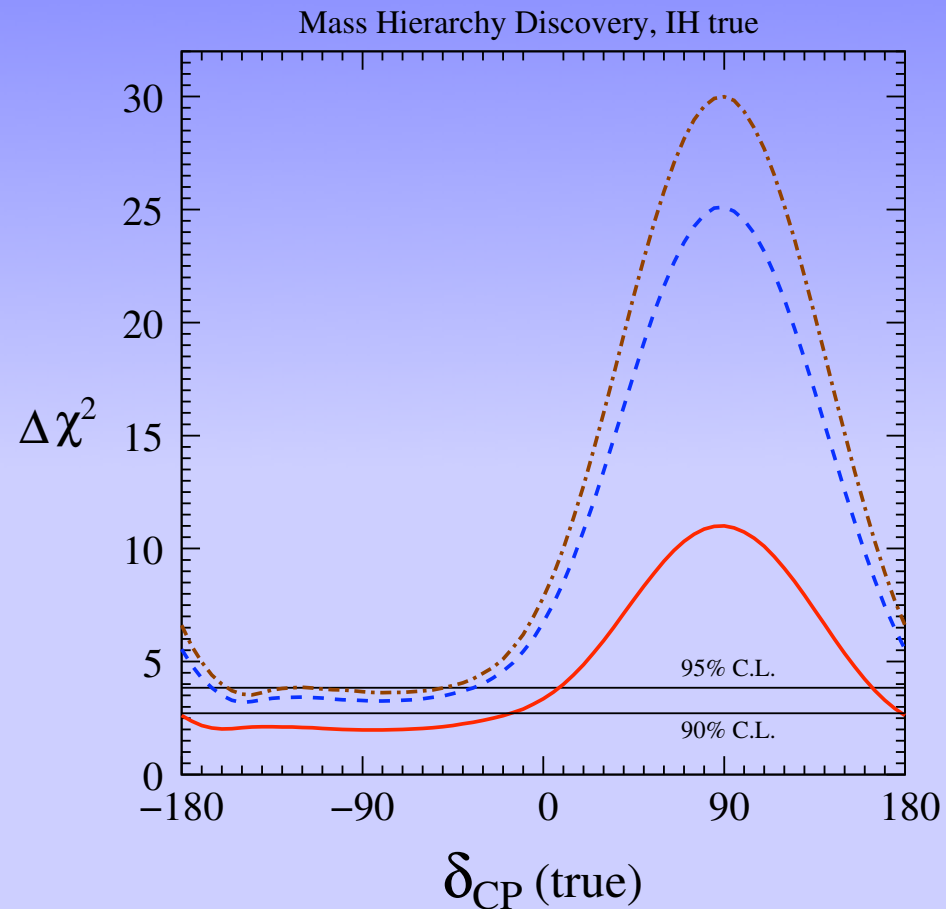
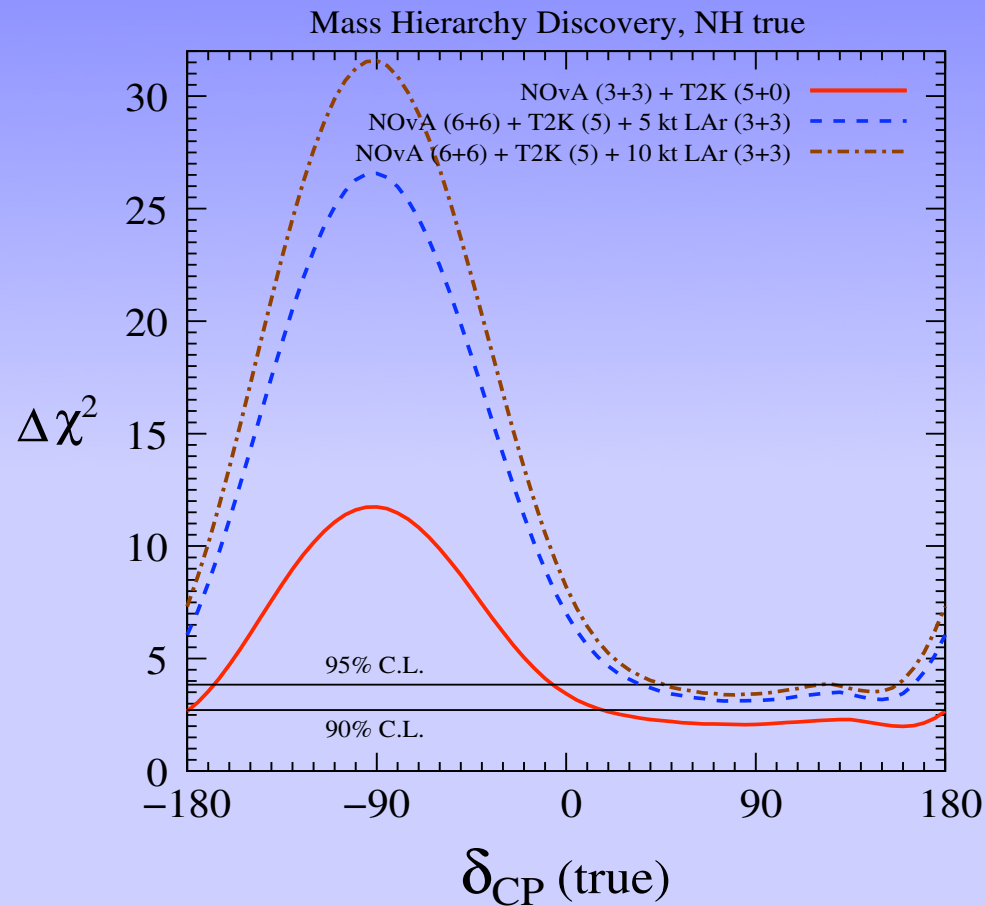
Present Generation Experiments: T2K and NOvA



Agarwalla, Prakash, Raut, Uma Sankar, arXiv:1208.3644

14kt NOvA is equivalent to 10 kt LArTPC

Add a small LArTPC in the NOvA Beamline

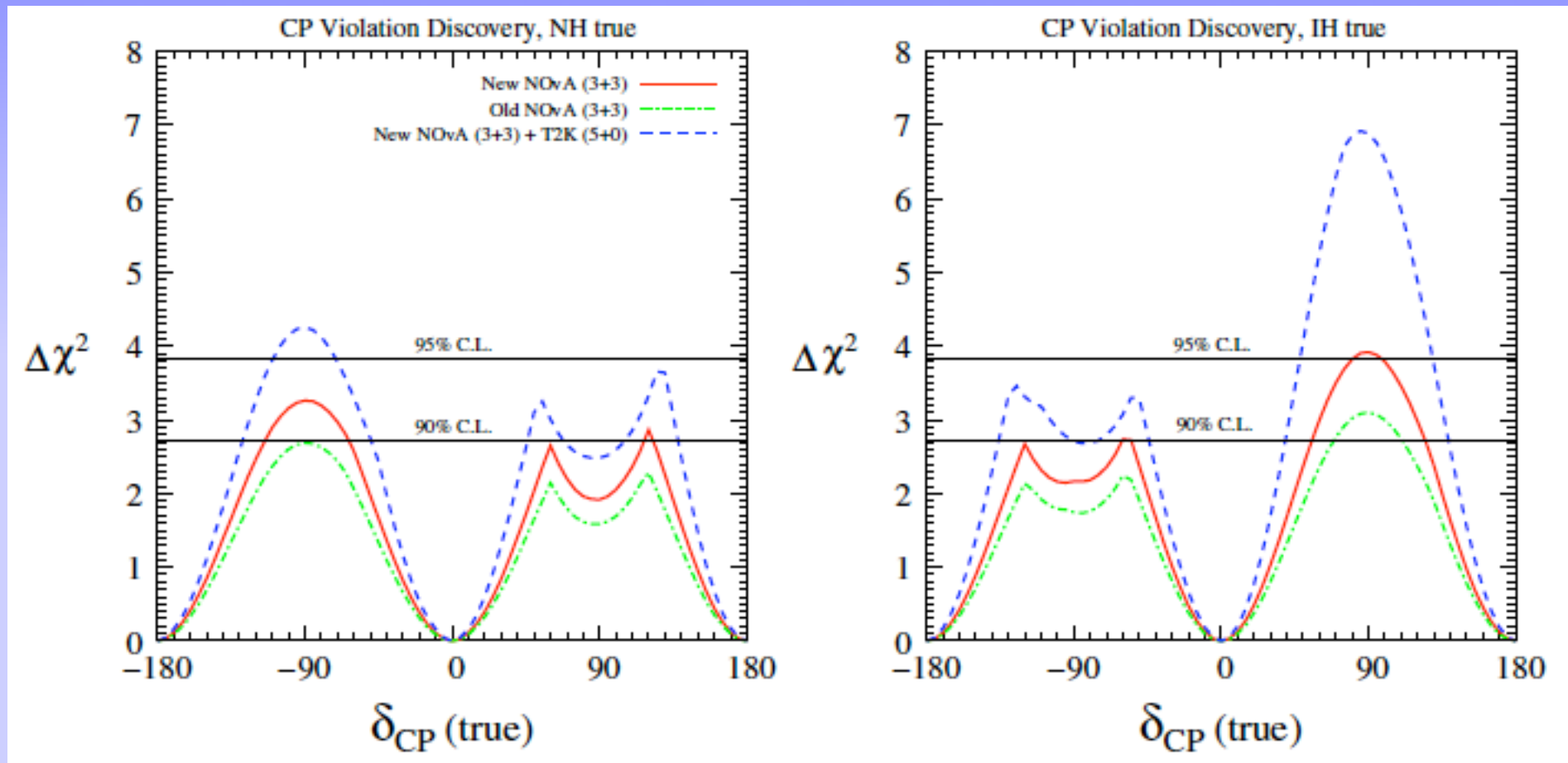


Agarwalla, Prakash, Raut, Uma Sankar, arXiv:1208.3644

Add a small LArTPC (5 to 10 kt) in the NOvA Beamline taking data simultaneously

100% CP coverage @ 90% C.L. and 64% CP coverage @ 95% C.L. w/ 5 kt LArTPC

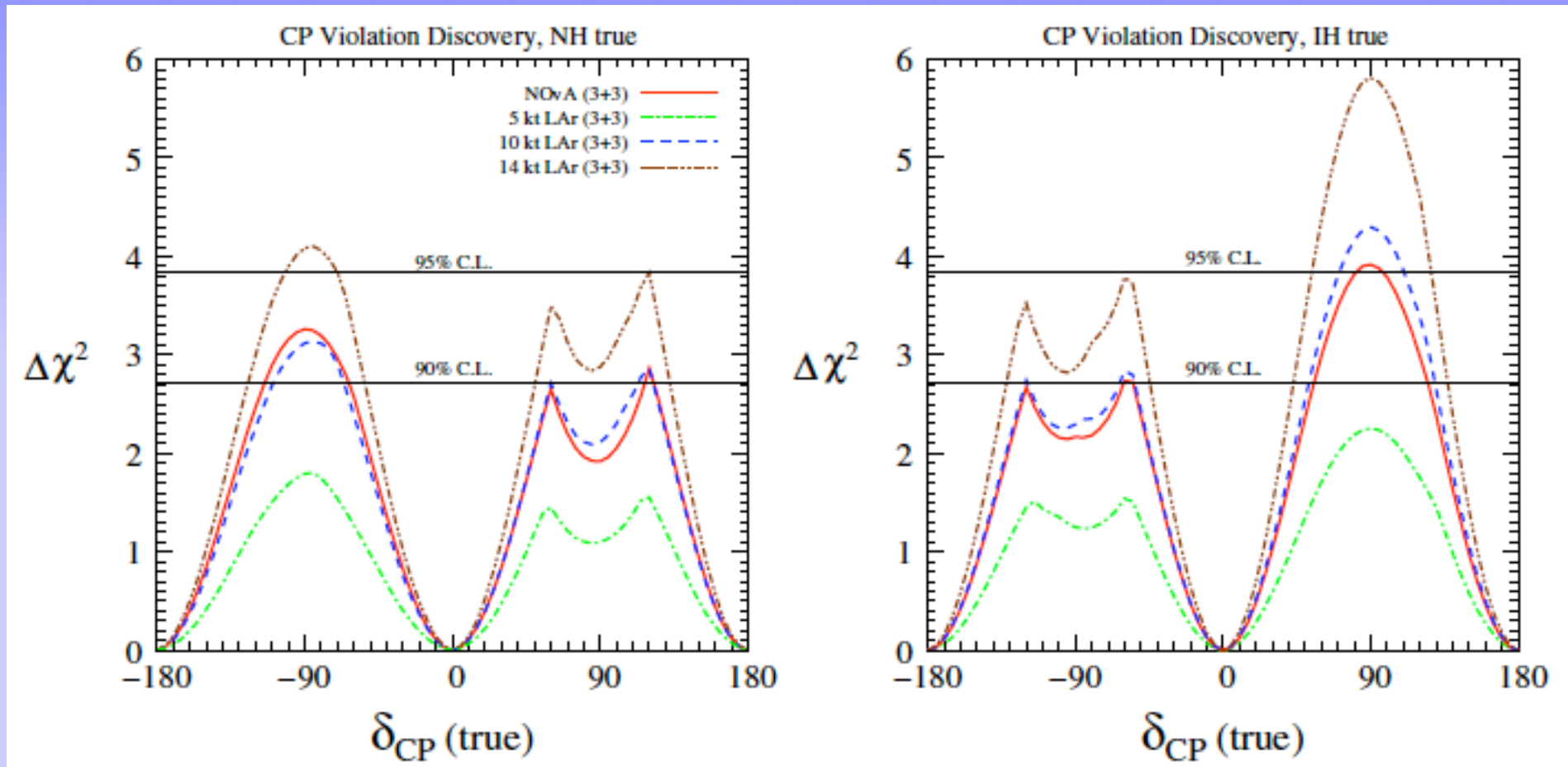
Present Generation Experiments: T2K and NOvA



Agarwalla, Prakash, Raut, Uma Sankar, arXiv:1208.3644

38% CP coverage @ 90% C.L. and 11% CP coverage @ 95% C.L. for CPV discovery

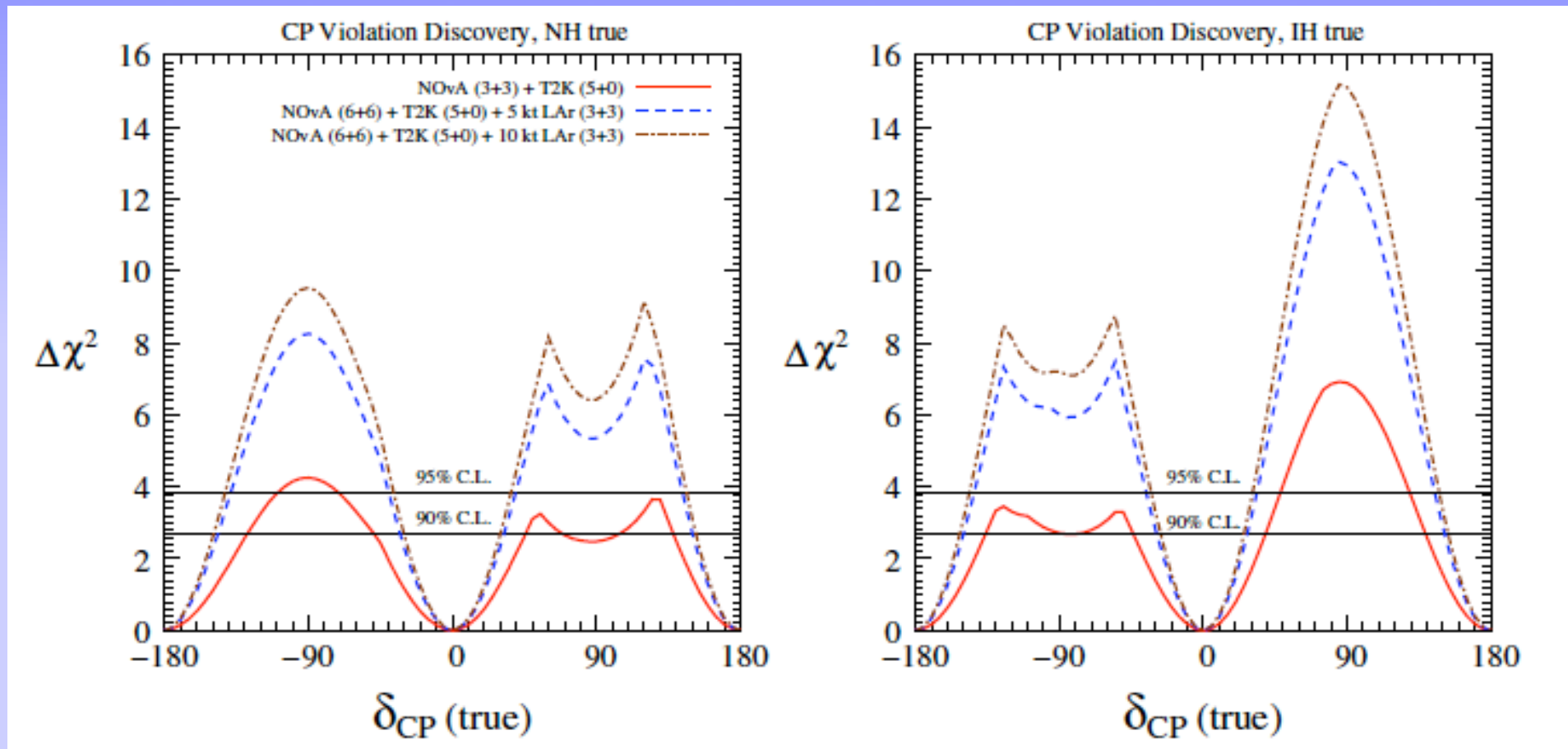
Present Generation Experiments: T2K and NOvA



Agarwalla, Prakash, Raut, Uma Sankar, arXiv:1208.3644

For CPV discovery: 14kt NOvA is equivalent to 10 kt LArTPC

Present Generation Experiments: T2K and NOvA



Agarwalla, Prakash, Raut, Uma Sankar, arXiv:1208.3644

64% CP coverage @ 90% C.L. and 56% CP coverage @ 95% C.L. w/ 5 kt LArTPC

Snapshot of MH & CPV discovery

$\sin^2 2\theta_{13}$	0.101	Fraction of $\delta_{CP}(\text{true})$			
		MH		CPV	
$\sin^2 \theta_{23}$	0.5	NH true	IH true	NH true	IH true
Setups					
NO ν A (3+3)		0.48 (0.43)	0.46 (0.41)	0.16 (0)	0.21(0.04)
NO ν A (3+3) + T2K (5+0)		0.55 (0.45)	0.54 (0.43)	0.38 (0.11)	0.49 (0.23)
NO ν A (6+6) + T2K (5+0) + 5 kt LArTPC (3+3)		1 (0.64)	1 (0.64)	0.64 (0.56)	0.68 (0.61)
NO ν A (6+6) + T2K (5+0) + 10 kt LArTPC (3+3)		1 (0.71)	1 (0.73)	0.67 (0.60)	0.71 (0.64)

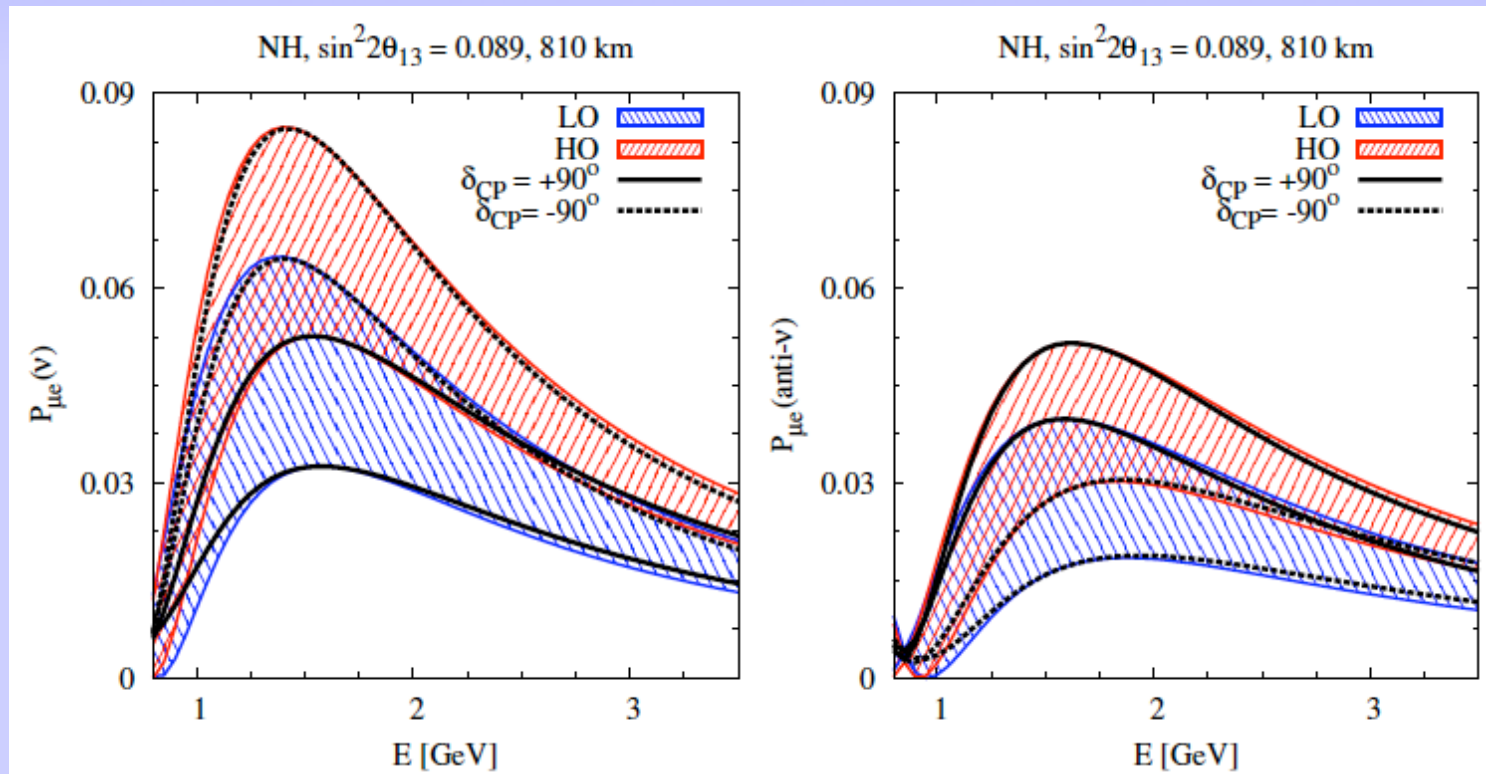
$\sin^2 2\theta_{13}$	0.089	Fraction of $\delta_{CP}(\text{true})$			
		MH		CPV	
$\sin^2 \theta_{23}$	0.413	NH true	IH true	NH true	IH true
Setups					
NO ν A (3+3)		0.39 (0.33)	0.37 (0.31)	0.2 (0.1)	0.22 (0.13)
NO ν A (3+3) + T2K (5+0)		0.41 (0.34)	0.39 (0.31)	0.28 (0.22)	0.3 (0.25)
NO ν A (6+6) + T2K (5+0) + 5 kt LArTPC (3+3)		0.78 (0.5)	0.89 (0.48)	0.68 (0.45)	0.71 (0.51)
NO ν A (6+6) + T2K (5+0) + 10 kt LArTPC (3+3)		1 (0.54)	1 (0.54)	0.7 (0.53)	0.73 (0.63)

Agarwalla, Prakash, Raut, Uma Sankar, arXiv:1208.3644

Non-maximal 2-3 mixing & Issue of Octant

Preliminary results of MINOS indicate that θ_{23} is not maximal

Global fits to world neutrino data suggest two nearly degenerate solutions for θ_{23} : one in the lower octant (LO: $\theta_{23} < 45$ degree) and the other in the higher octant (HO: $\theta_{23} > 45$ degree)

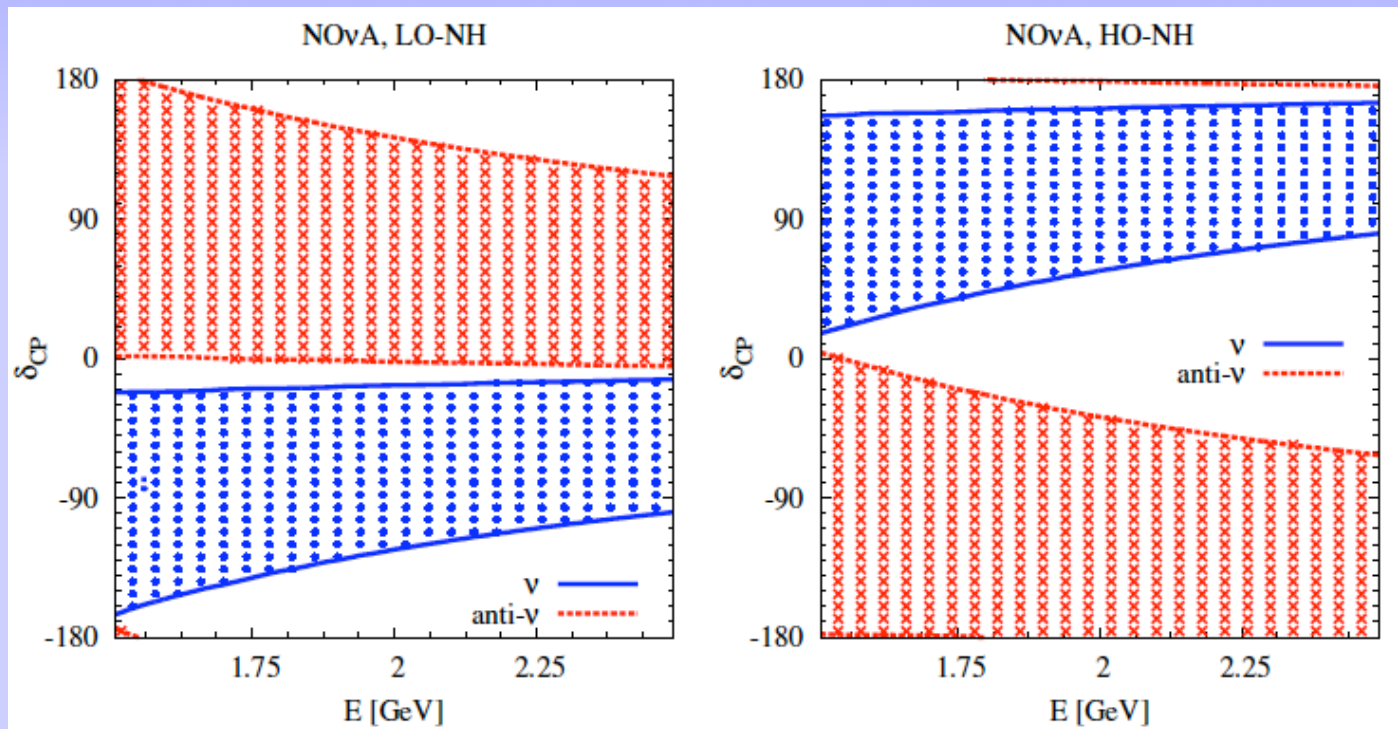


Agarwalla, Prakash, Uma Sankar, arXiv:1301.2574

Octant- δ_{CP} degeneracy

$$P_{\mu e}(LO, \delta_{CP}^{LO}) = P_{\mu e}(HO, \delta_{CP}^{HO})$$

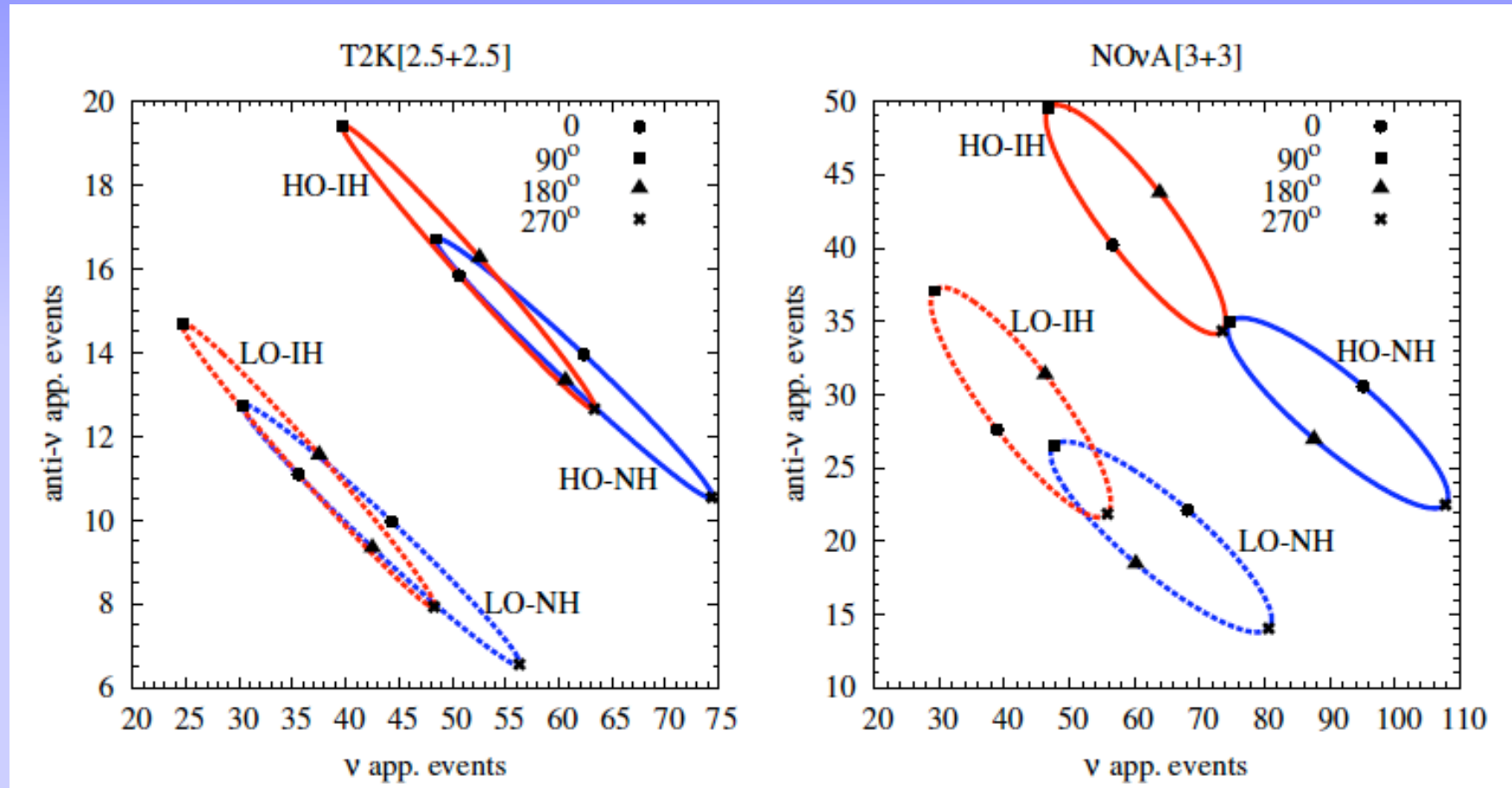
For $L=810$ km and $E=2$ GeV, we get for NH and ν $\cos(\hat{\Delta} + \delta_{CP}^{LO}) - \cos(\hat{\Delta} + \delta_{CP}^{HO}) = 1.7$



Agarwalla, Prakash, Uma Sankar, arXiv:1301.2574

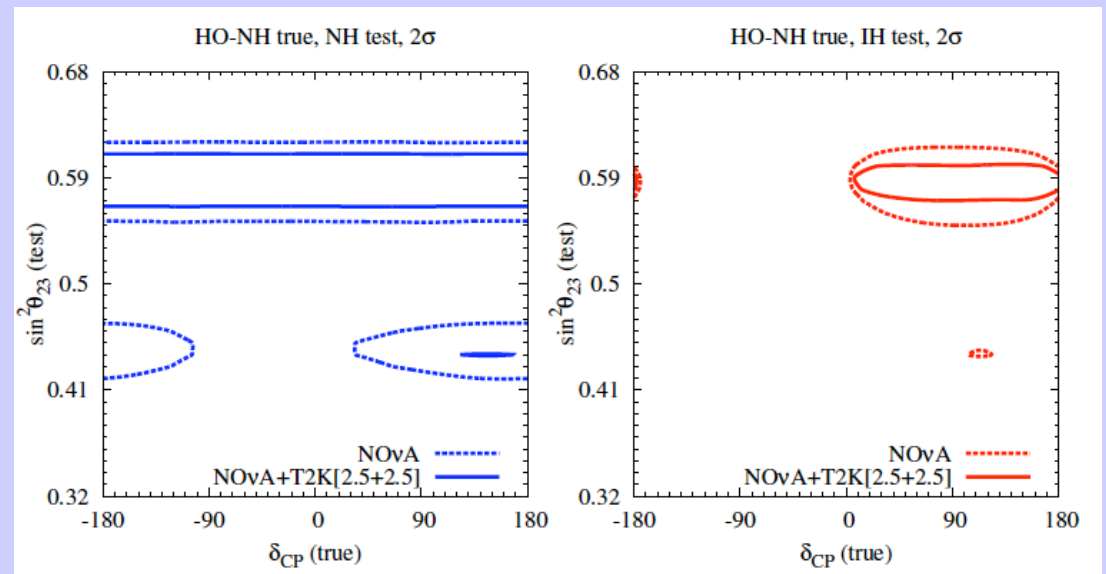
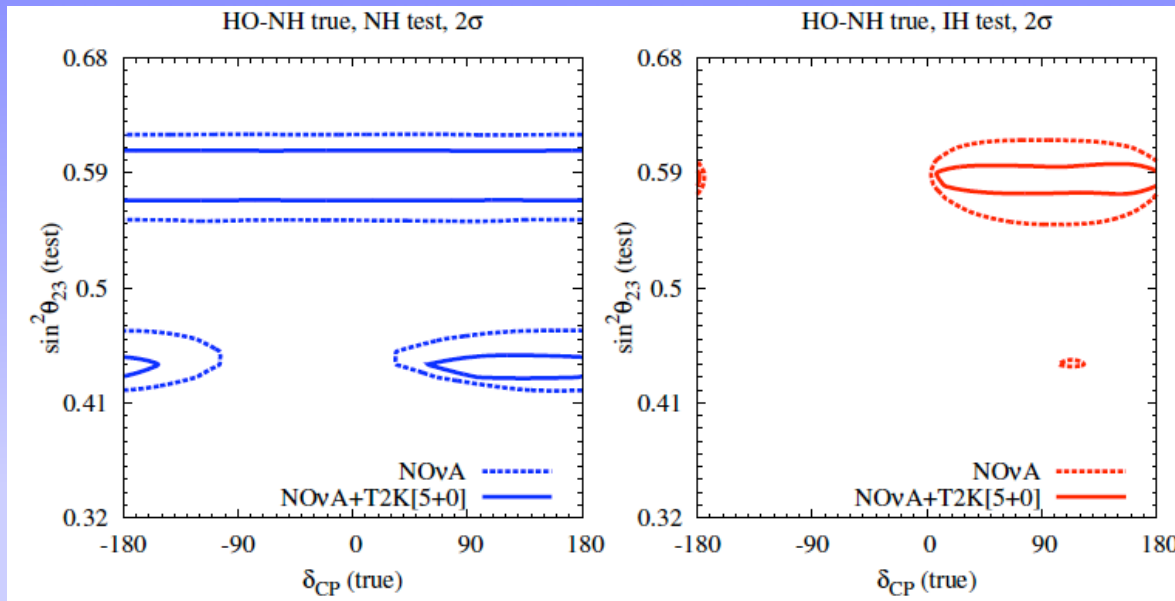
$P_{\mu e}(LO, -116^\circ \leq \delta_{CP} \leq -26^\circ)$ is degenerate with $P_{\mu e}(HO, 64^\circ \leq \delta_{CP} \leq 161^\circ)$

Octant-Hierarchy combinations: T2K and NOvA



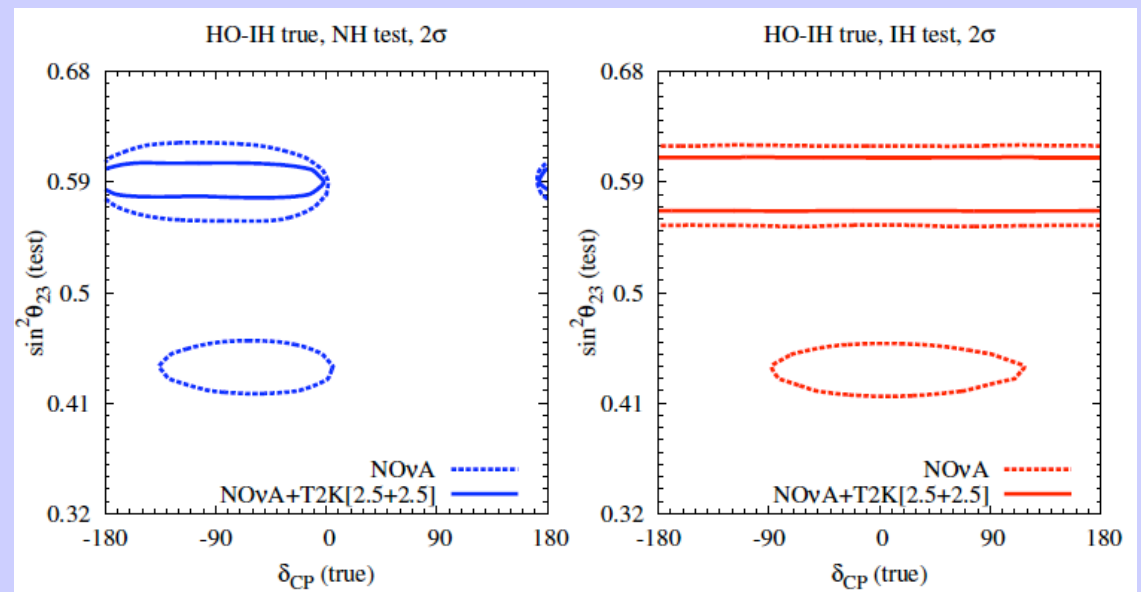
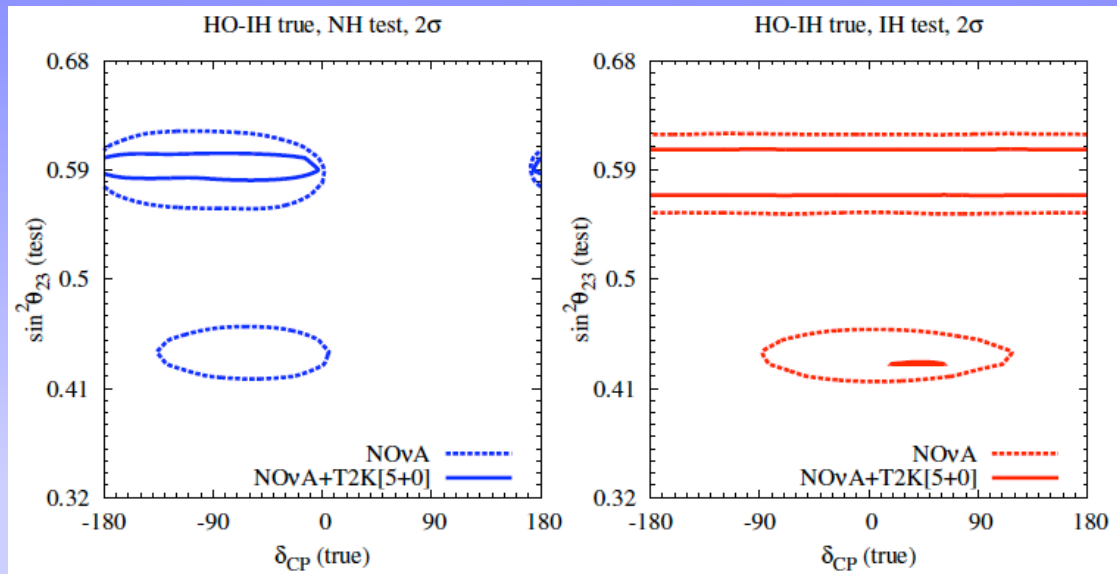
Agarwalla, Prakash, Uma Sankar, arXiv:1301.2574

Octant Resolution: T2K and NOvA



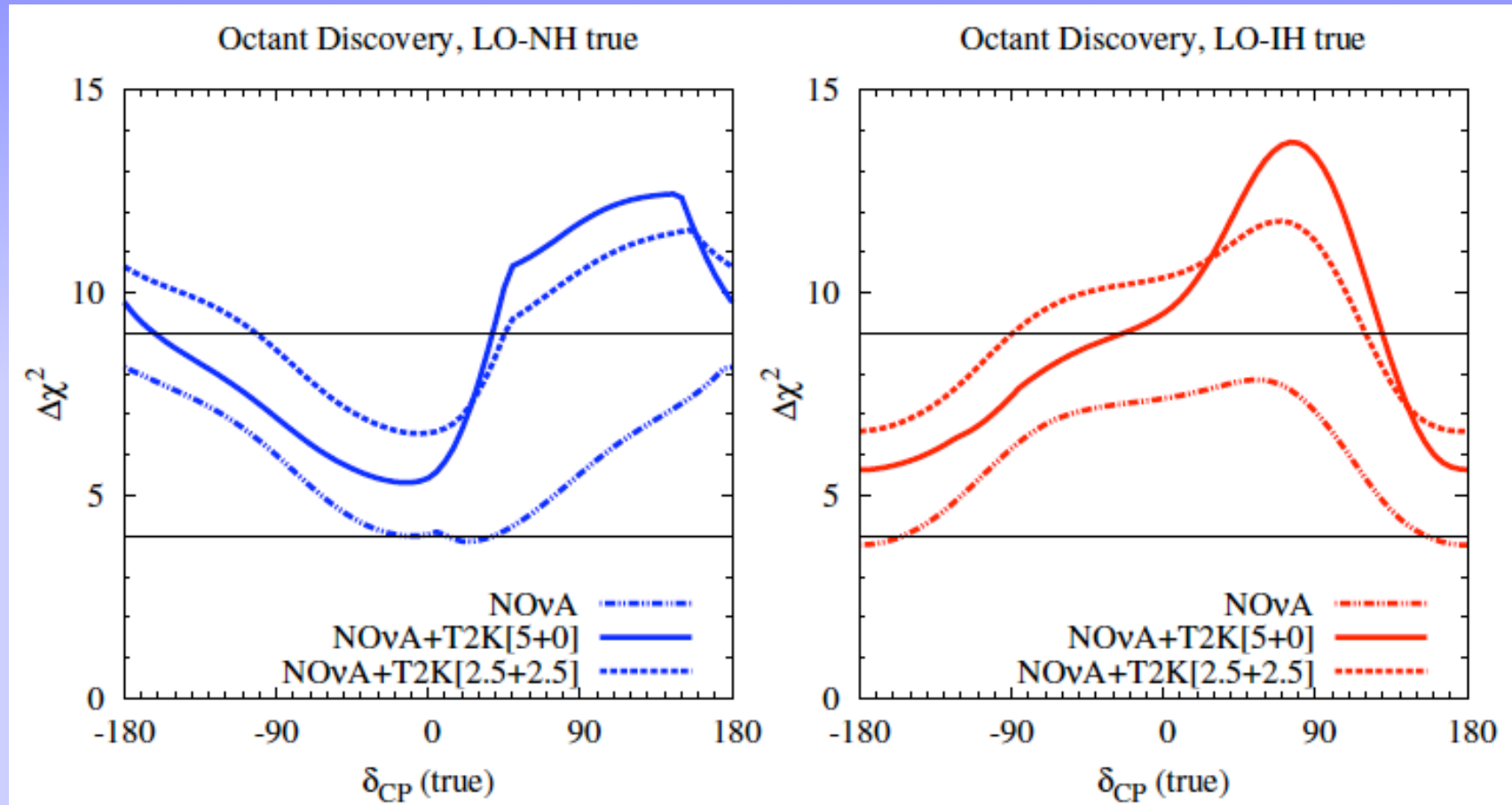
Agarwalla, Prakash, Uma Sankar, arXiv:1301.2574

Octant Resolution: T2K and NOvA



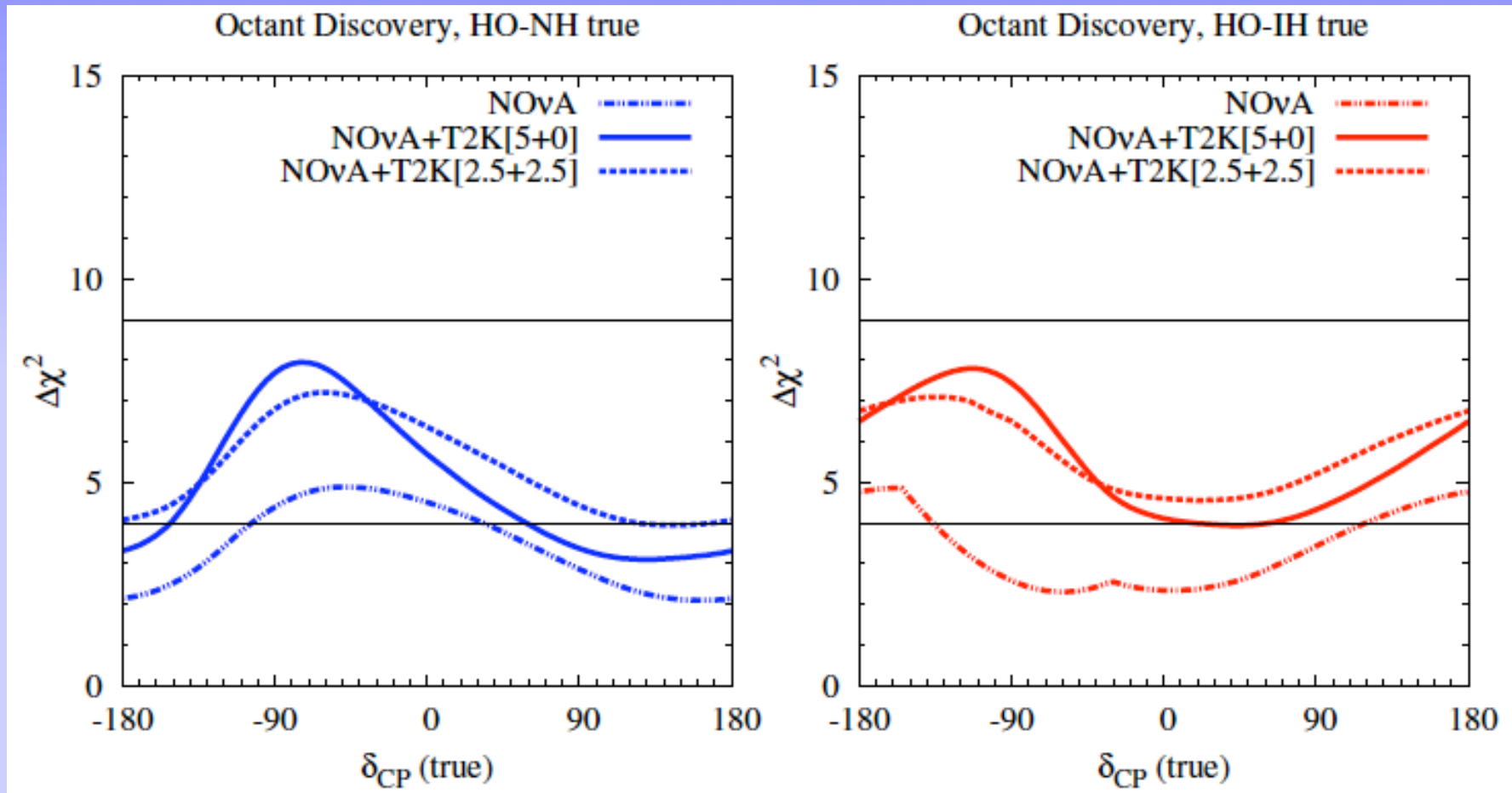
Agarwalla, Prakash, Uma Sankar, arXiv:1301.2574

Octant Discovery: T2K and NOvA



Agarwalla, Prakash, Uma Sankar, arXiv:1301.2574

Octant Discovery: T2K and NOvA



Agarwalla, Prakash, Uma Sankar, arXiv:1301.2574

Octant Discovery: T2K and NOvA

- If LO is the true octant, then NOvA can rule out HO at 2σ C.L., irrespective of the hierarchy and CP phase
- Addition of T2K data improves the octant sensitivity
- If HO and normal hierarchy are the true choices, then the combined data from NOvA & T2K with its designed five years run in neutrino mode, is incapable of a 2σ resolution of the octant for all CP phases
- A 2σ resolution of the octant, for all combinations of neutrino parameters, becomes possible if T2K has balanced neutrino and anti-neutrino runs of 2.5 years each

Thank you!

Platinum Channel ($P_{\mu e}$)

The appearance probability ($\nu_\mu \rightarrow \nu_e$) in matter, upto second order in the small parameters $\alpha \equiv \Delta m_{21}^2 / \Delta m_{31}^2$ and $\sin 2\theta_{13}$,

$$\begin{aligned}
 P_{\mu e} \simeq & \frac{\sin^2 2\theta_{13} \sin^2 \theta_{23}}{0.05} \frac{\sin^2[(1 - \hat{A})\Delta]}{(1 - \hat{A})^2} \Rightarrow \theta_{13} \text{ Driven} \\
 & - \frac{\alpha \sin 2\theta_{13} \xi \sin \delta_{CP}}{0.0096} \sin(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1 - \hat{A})\Delta]}{(1 - \hat{A})} \Rightarrow \text{CP odd} \\
 & + \alpha \sin 2\theta_{13} \xi \cos \delta_{CP} \cos(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1 - \hat{A})\Delta]}{(1 - \hat{A})} \Rightarrow \text{CP even} \\
 & + \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}; \Rightarrow \text{Solar Term}
 \end{aligned}$$

where $\Delta \equiv \Delta m_{31}^2 L / (4E)$, $\xi \equiv \cos \theta_{13} \sin 2\theta_{21} \sin 2\theta_{23}$,
and $\hat{A} \equiv \pm(2\sqrt{2}G_F n_e E) / \Delta m_{31}^2$

Cervera et al., hep-ph/0002108
Freund et al., hep-ph/0105071