

The Lepton Flavor Violation road to New Physics

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IMHEP 2019
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CONSEJO SUPERIOR DE INVESTIGACIONES CIENTÍFICAS



Introduction

**Before the LHC started operating we all
hoped for **great discoveries**...**

A dense tropical rainforest with sunlight filtering through the canopy. The scene is filled with various types of green plants, including palm trees and broad-leafed species. The lighting is bright, creating a high-contrast scene with deep shadows and bright highlights where the sun hits the leaves.

Microscopic
black holes

Extra dimensions

Supersymmetry

Compositeness

LHC expectations

LHC results...

**125 GeV
palm tree**



Introduction

Do we have a good reason to go **Beyond the Standard Model?**

Introduction

Do we have a good reason to go **Beyond the Standard Model?**



Neutrinos!

The lepton sector is still
to be understood!

Neutrinos and the lepton sector

Dear radioactive Ladies and Gentlemen...

Physikalisches Institut
der Eidg. Technischen Hochschule
Zürich

Zürich, 4. Dez. 1930
Gloriastrasse

Liebe Radioaktive Damen und Herren,

Wie der Ueberbringer dieser Zeilen, den ich halbvollst anzuhören bitte, Ihnen des näheren auseinandersetzen wird, bin ich angesichts der "falschen" Statistik der N- und Li-6 Kerne, sowie des kontinuierlichen beta-Spektrums auf einen verzweifelten Ausweg verfallen um den "Wechselsatz" (1) der Statistik und den Energiesatz zu retten. Nämlich die Möglichkeit, es könnten elektrisch neutrale Teilchen, die ich Neutronen nennen will, in den Kernen existieren, welche den Spin $1/2$ haben und das Ausschliessungsprinzip befolgen und sich von Lichtquanten ausserdem noch dadurch unterscheiden, dass sie nicht mit Lichtgeschwindigkeit laufen. Die Masse der Neutronen dürfte von derselben Grössenordnung wie die Elektronenmasse sein und jedenfalls nicht grösser als $0,01$ Protonenmasse.- Das kontinuierliche beta-Spektrum wäre dann verständlich unter der Annahme, dass beim beta-Zerfall mit dem Elektron jeweils noch ein Neutron emittiert wird, derart, dass die Summe der Energien von Neutron und Elektron konstant ist.



December 4th, 1930

Letter to his colleagues in Tübingen

1930

Pauli's neutrino hypothesis

Open questions

What is the origin of neutrinos masses?

Are they Dirac or Majorana?

What is the absolute scale of neutrino masses?

What is the mass ordering?

Are there more than three neutrinos? Maybe sterile?

Is there CP violation in the lepton sector?

Lepton flavor violation

In the **Standard Model**, three copies of the leptonic SU(2) doublet are introduced

$$\begin{pmatrix} \nu_e \\ e^- \end{pmatrix} \quad \begin{pmatrix} \nu_\mu \\ \mu^- \end{pmatrix} \quad \begin{pmatrix} \nu_\tau \\ \tau^- \end{pmatrix} \quad \begin{array}{c} \updownarrow \\ \text{Gauge and} \\ \text{Yukawa} \\ \text{interactions} \end{array}$$

←—————→

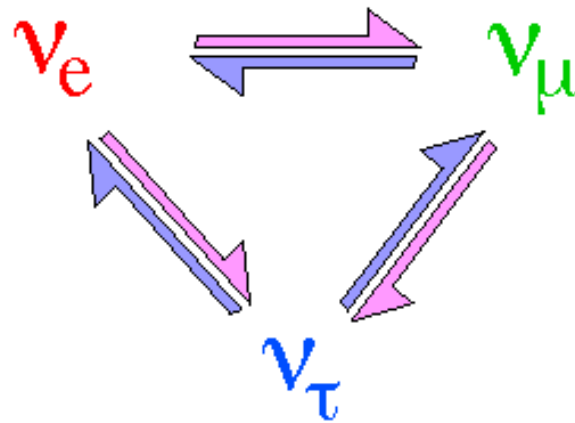
X

Is **lepton flavor** a conserved quantity?

Neutrino oscillations: LFV

We already know the answer: **NO**

Neutrino flavor oscillations: flavor violating process!



Then... is this process a sign of BSM physics?

Yes it is!



What about cLFV?

In conclusion, lepton flavor is **not** conserved: there is **lepton flavor violation (LFV)**

However... what about **charged lepton flavor violation (cLFV)**?

$$\mu^- \rightarrow e^- \gamma$$

$$h \rightarrow \mu^- \tau^+$$

$$\tau^- \rightarrow \mu^- \mu^+ \mu^-$$

$$\pi^0 \rightarrow e^- \mu^+$$

$$K_L^0 \rightarrow \pi^0 e^- \mu^+$$

...

Never observed...

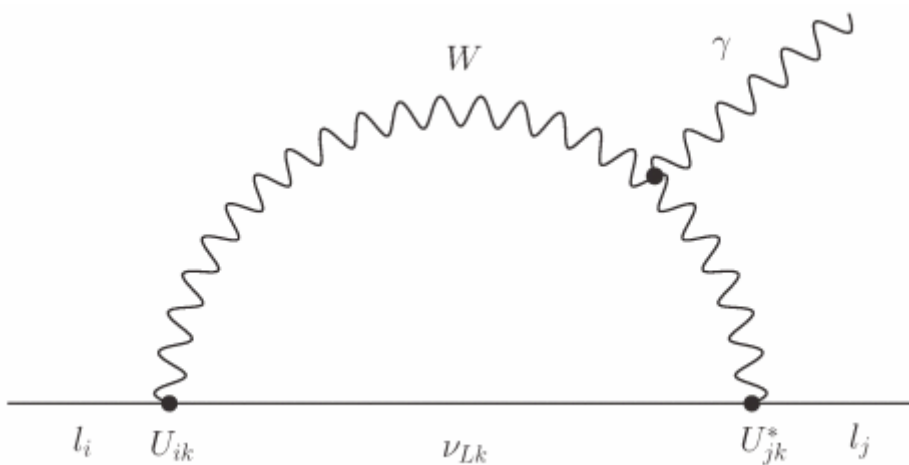
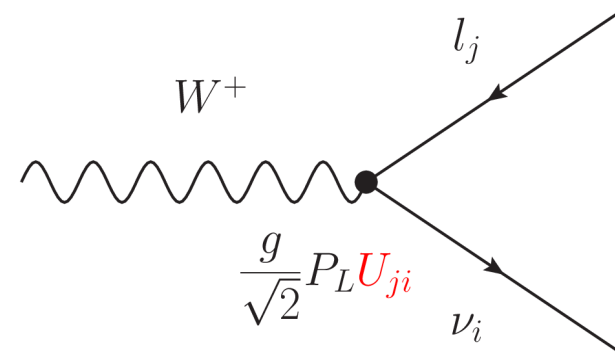
What about cLFV?

SM + Dirac neutrino masses

$$-\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} U_{ji} \bar{l}_j \gamma^\mu P_L \nu_i W_\mu^- + \text{c.c.}$$

U : lepton mixing matrix

[analog of the CKM matrix in the lepton sector]



$$\text{Br}(\mu \rightarrow e \gamma) = \frac{3\alpha}{32\pi} \left| \sum_k U_{ek} U_{\mu k}^* \frac{m_{\nu k}^2}{m_W^2} \right|^2 \lesssim 10^{-54}$$

Since neutrino masses are the **only source** of LFV, all cLFV amplitudes are strongly suppressed (in fact, **GIM** suppressed)

Why do we care about LFV?

The observation of **cLFV** would be a clear signal of (non-trivial) physics beyond the Standard Model

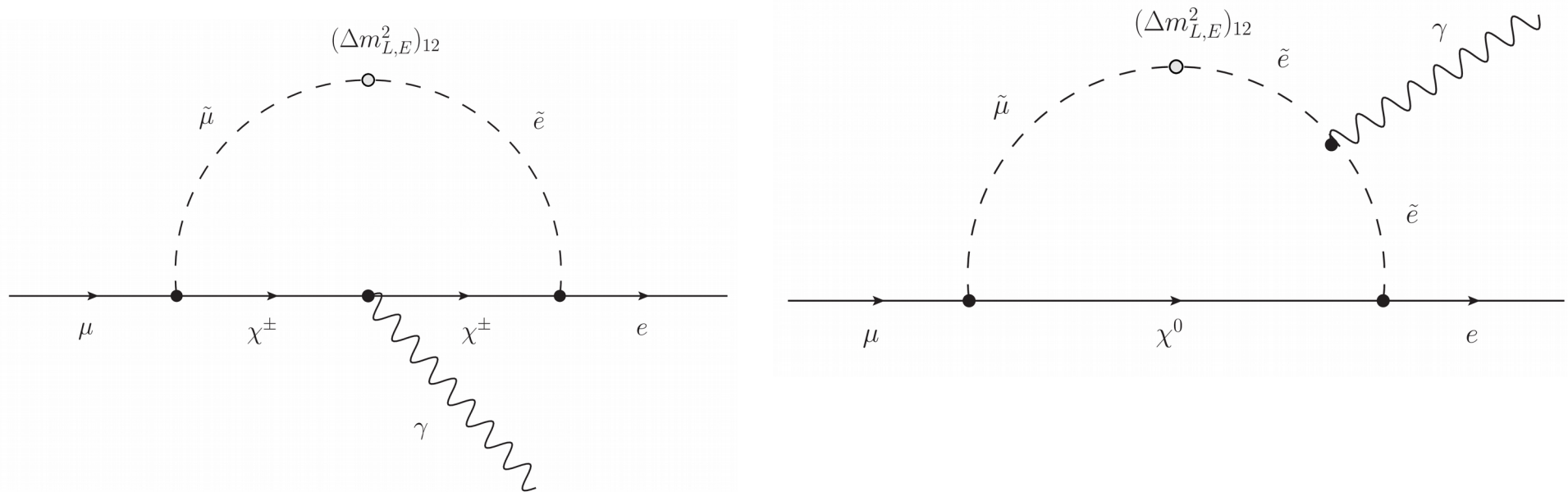
In fact, most **BSM models** predict **large cLFV rates**

We can probe **very high energy scales!**

$$\mathcal{O} = \frac{c_{e\mu}}{\Lambda^2} \bar{\mu} e \bar{e} e \quad \Rightarrow \quad \frac{\Lambda}{\sqrt{c_{e\mu}}} \gtrsim 100 \text{ TeV}$$

Why do we care about LFV?

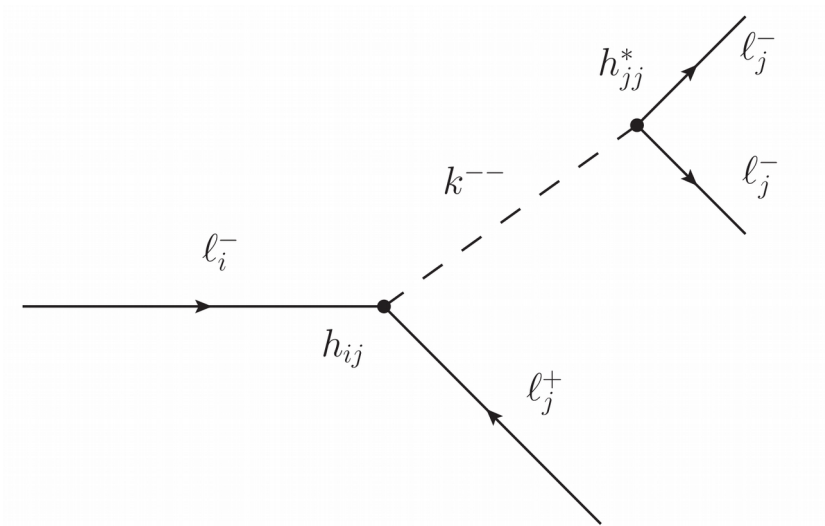
Example 1: Supersymmetric models



Sleptons: a whole new sector coupled to the **SM leptons**
Strong constraints on the **off-diagonal** soft terms

Why do we care about LFV?

Example 2: Babu-Zee model



$$\text{BR} \sim \left| \frac{h_{ij} h_{jj}^*}{m_k^2} \right|^2$$

Small **off-diagonal h couplings** and/or **heavy k's** are required

Experimental projects

LFV Process	Present Bound	Future Sensitivity
$\mu \rightarrow e\gamma$	4.2×10^{-13}	6×10^{-14} (MEG)
$\tau \rightarrow e\gamma$	3.3×10^{-8}	$\sim 10^{-8} - 10^{-9}$ (B factories)
$\tau \rightarrow \mu\gamma$	4.4×10^{-8}	$\sim 10^{-8} - 10^{-9}$ (B factories)
$\mu \rightarrow 3e$	1.0×10^{-12}	$\sim 10^{-16}$ (Mu3e)
$\tau \rightarrow 3e$	2.7×10^{-8}	$\sim 10^{-9} - 10^{-10}$ (B factories)
$\tau \rightarrow 3\mu$	2.1×10^{-8}	$\sim 10^{-9} - 10^{-10}$ (B factories)
$\mu^-, \text{Au} \rightarrow e^-, \text{Au}$	7.0×10^{-13}	—
$\mu^-, \text{SiC} \rightarrow e^-, \text{SiC}$	—	2×10^{-14} (DeeMe)
$\mu^-, \text{Al} \rightarrow e^-, \text{Al}$	—	$10^{-15} - 10^{-17}$ (COMET)
$\mu^-, \text{Ti} \rightarrow e^-, \text{Ti}$	4.3×10^{-12}	$10^{-17} - 10^{-18}$ (Mu2e)
		$\sim 10^{-18}$ (PRISM/PRIME)

Experimental projects

History of $\mu \rightarrow e\gamma$, $\mu N \rightarrow eN$, and $\mu \rightarrow 3e$

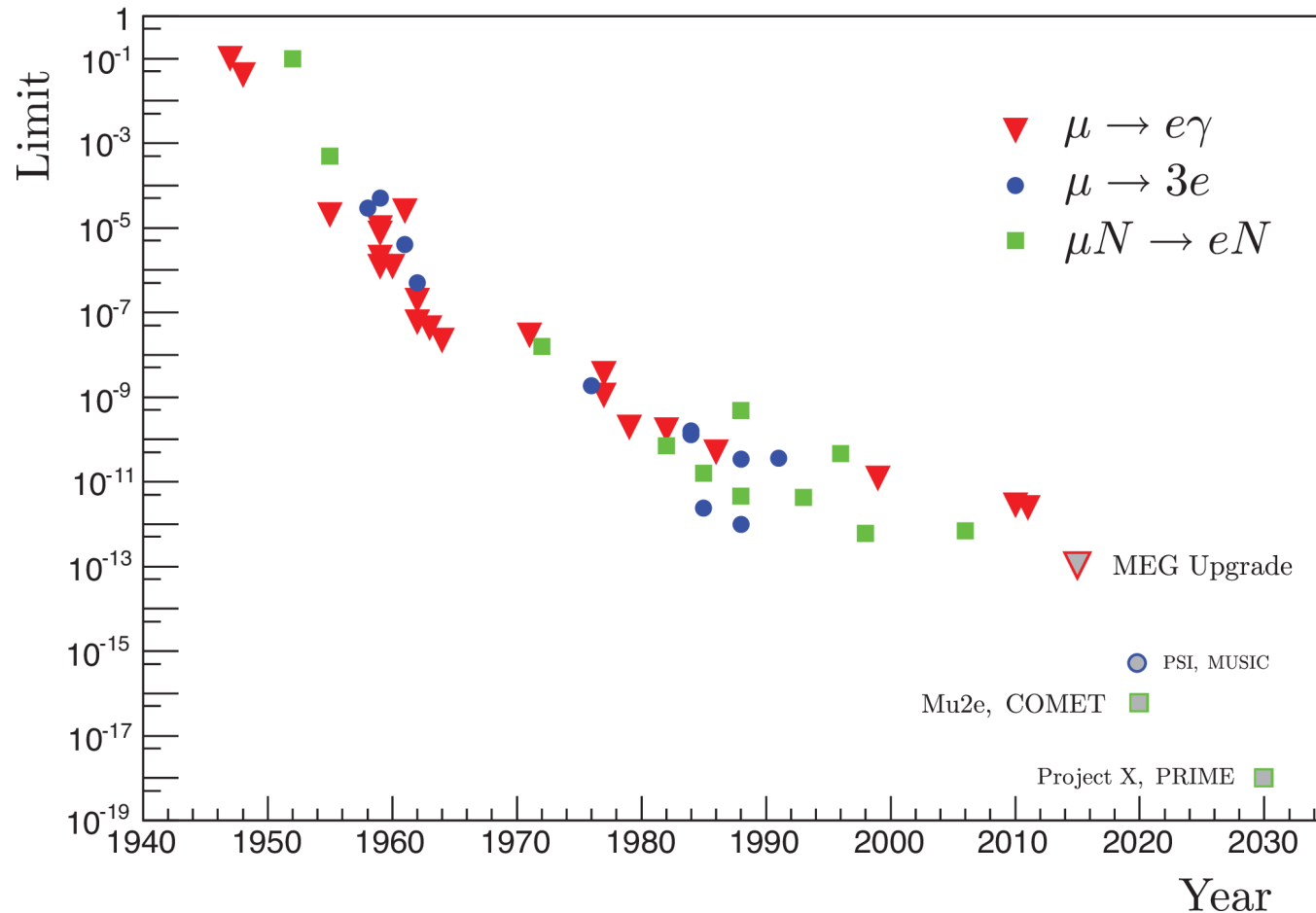
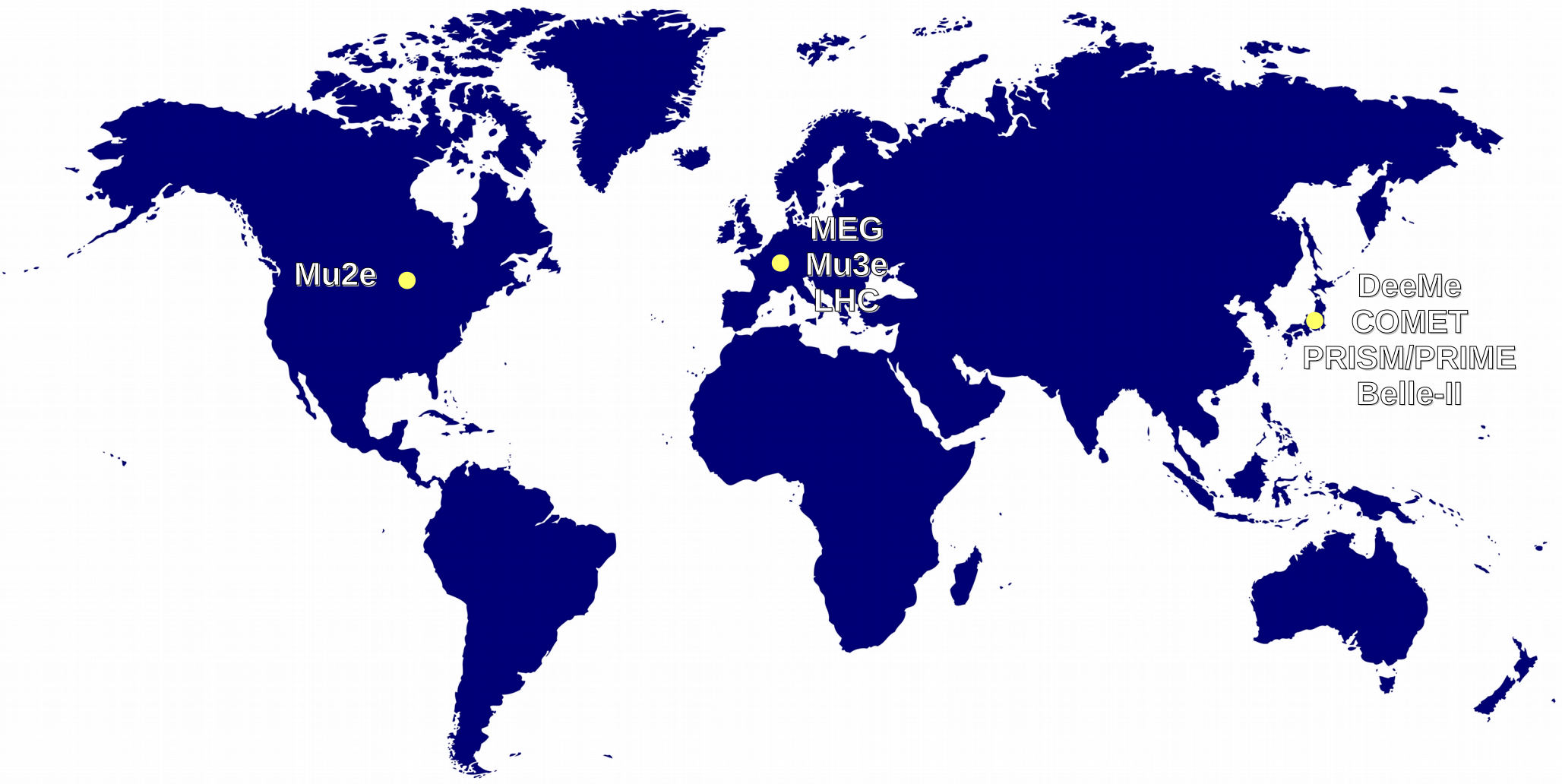


Figure taken from Bernstein & Cooper [arXiv:1307.5787]

Experimental projects



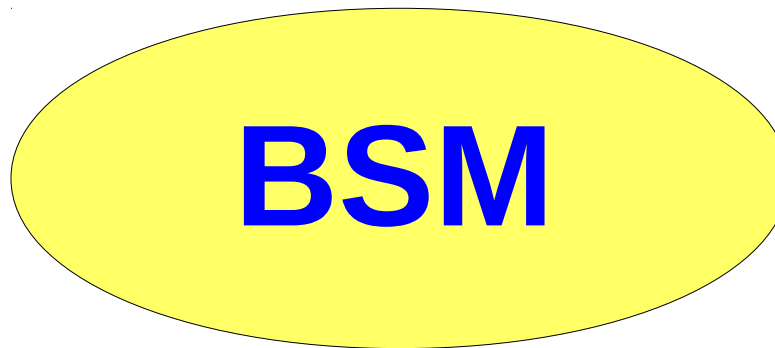
LFV : Where to look for?

$$l_i \rightarrow l_j \gamma$$

$$l_i \rightarrow 3 l_j$$

$$l_i \rightarrow l_j l_k l_k$$

$\mu - e$
conversion in nuclei



LFV at colliders

$$M \rightarrow l_i l_j$$

LFV : Where to look for?

Everywhere!

ratio	LHT	MSSM (dipole)	MSSM (Higgs)	SM4
$\frac{\text{Br}(\mu^- \rightarrow e^- e^+ e^-)}{\text{Br}(\mu \rightarrow e \gamma)}$	0.02... 1	$\sim 6 \cdot 10^{-3}$	$\sim 6 \cdot 10^{-3}$	0.06... 2.2
$\frac{\text{Br}(\tau^- \rightarrow e^- e^+ e^-)}{\text{Br}(\tau \rightarrow e \gamma)}$	0.04... 0.4	$\sim 1 \cdot 10^{-2}$	$\sim 1 \cdot 10^{-2}$	0.07... 2.2
$\frac{\text{Br}(\tau^- \rightarrow \mu^- \mu^+ \mu^-)}{\text{Br}(\tau \rightarrow \mu \gamma)}$	0.04... 0.4	$\sim 2 \cdot 10^{-3}$	0.06... 0.1	0.06... 2.2
$\frac{\text{Br}(\tau^- \rightarrow e^- \mu^+ \mu^-)}{\text{Br}(\tau \rightarrow e \gamma)}$	0.04... 0.3	$\sim 2 \cdot 10^{-3}$	0.02... 0.04	0.03... 1.3
$\frac{\text{Br}(\tau^- \rightarrow \mu^- e^+ e^-)}{\text{Br}(\tau \rightarrow \mu \gamma)}$	0.04... 0.3	$\sim 1 \cdot 10^{-2}$	$\sim 1 \cdot 10^{-2}$	0.04... 1.4
$\frac{\text{Br}(\tau^- \rightarrow e^- e^+ e^-)}{\text{Br}(\tau^- \rightarrow e^- \mu^+ \mu^-)}$	0.8... 2	~ 5	0.3... 0.5	1.5... 2.3
$\frac{\text{Br}(\tau^- \rightarrow \mu^- \mu^+ \mu^-)}{\text{Br}(\tau^- \rightarrow \mu^- e^+ e^-)}$	0.7... 1.6	~ 0.2	5... 10	1.4... 1.7
$\frac{\text{R}(\mu \text{Ti} \rightarrow e \text{Ti})}{\text{Br}(\mu \rightarrow e \gamma)}$	$10^{-3} \dots 10^2$	$\sim 5 \cdot 10^{-3}$	0.08... 0.15	$10^{-12} \dots 26$

Table taken from Buras et al [arXiv:1006.5356]

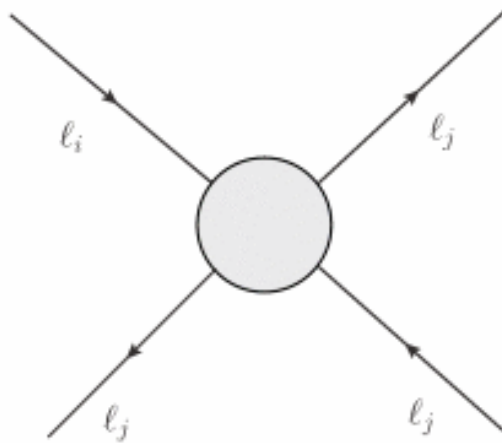
$$l_i \rightarrow 3 l_j \text{ VS } l_i \rightarrow l_j \gamma$$

What contribution dominates $l_i \rightarrow 3 l_j$?

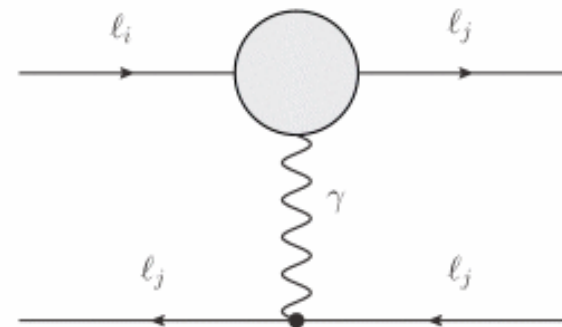
In many models of interest: **Photonic dipole contributions**

Most popular example: MSSM

[Hisano et al 1996; Arganda, Herrero 2006]



\approx



Dipole dominance

$$\frac{BR(l_i \rightarrow 3 l_j)}{BR(l_i \rightarrow l_j \gamma)} = \frac{\alpha}{3\pi} \left(\log \frac{m_{l_i}^2}{m_{l_j}^2} - \frac{11}{4} \right) \Rightarrow BR(l_i \rightarrow l_j \gamma) \gg BR(l_i \rightarrow 3 l_j)$$

The LFV program

In order to unravel the **physics behind LFV** (and perhaps neutrino masses!) we must:

- **Search for LFV in as many observables as possible:** they might have information about different sectors of the theory
- **Study the relations among different observables** (ratios, correlations, hierarchies...)
- **Understand the origin of such relations:** what is the underlying physics?

Outline of the talk

- Introduction: Lepton Flavor Violation
- Selected topics
 - LFV in low-scale seesaw models
 - LFV in B-meson decays
 - Master Majorana parametrization
- Final remarks



Chuck Norris fact of the day

*Chuck Norris counted to
infinity. Twice.*



LFV in low-scale seesaw models

**With Asmaa Abada, Manuel E. Krauss, Werner
Porod, Florian Staub and Cédric Weiland**

[PRD 90 \(2014\) 013008 \[arXiv:1312.5318\]](#)

[JHEP 1411 \(2014\) 048 \[arXiv:1408.0138\]](#)

Low-scale seesaw models

[Mohapatra, Valle, 1986]

The Inverse Seesaw

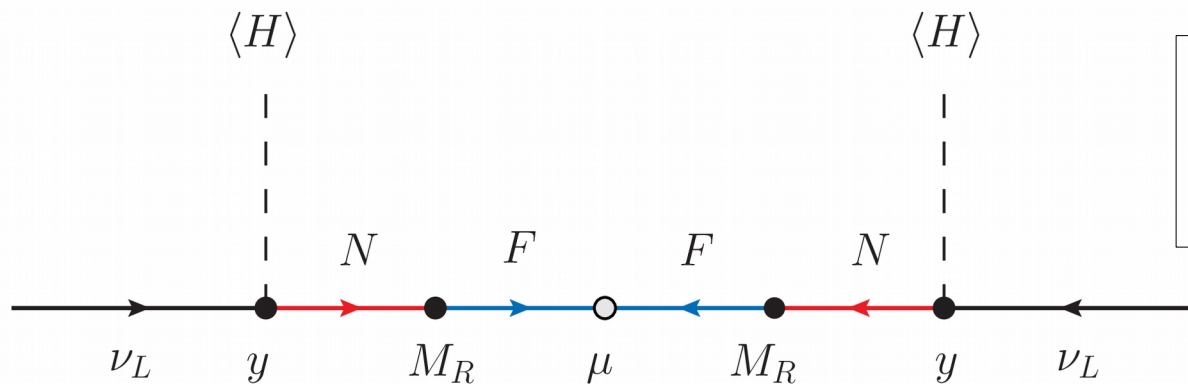
$$-\mathcal{L}_{IS} \supset y_{ij} \bar{N}_i L_j \tilde{H} + M_{Rij} \bar{N}_i F_j + \frac{1}{2} \mu_{ij} \bar{F}_i^c F_j$$

singlet states: $\#N = \#F = 3$

9x9 mass matrix



$$\mathcal{M} = \begin{pmatrix} 0 & \frac{1}{\sqrt{2}} y^T v & 0 \\ \frac{1}{\sqrt{2}} y v & 0 & M_R \\ 0 & M_R^T & \mu \end{pmatrix}$$



$$m \simeq \frac{v^2}{2} y^T (M_R^T)^{-1} \mu M_R^{-1} y$$

[if $\mu \ll yv \ll M_R$]

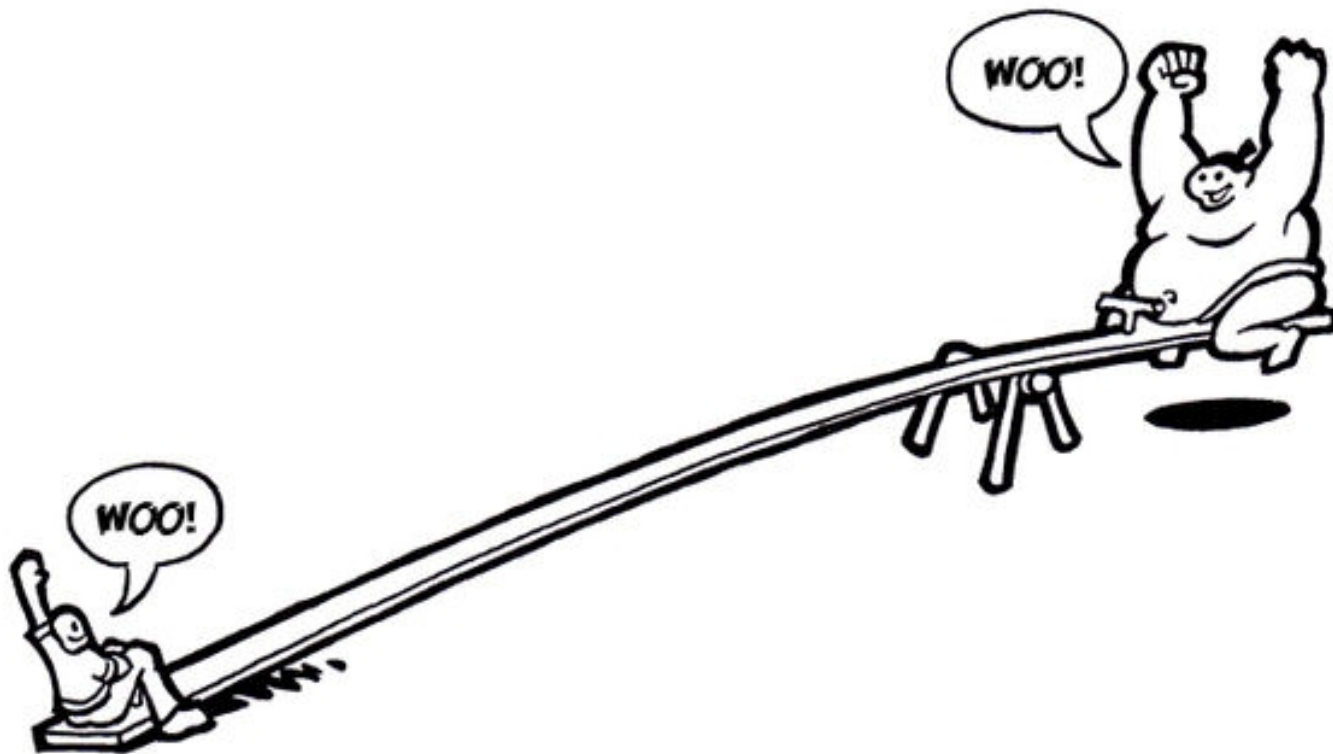
Standard vs Inverse Seesaw

Standard Seesaw



Standard vs Inverse Seesaw

Inverse Seesaw

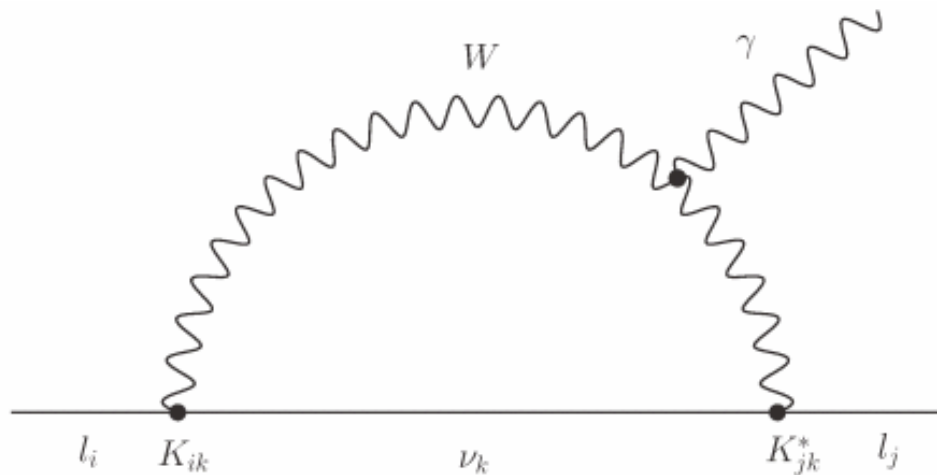


Penguins in the inverse seesaw

[Ilakovac, Pilaftsis, 1995; Deppisch, Valle, 2005]

$$\text{Br}(\mu \rightarrow e\gamma) = \frac{\alpha_W^3 s_W^2 m_\mu^5}{256\pi^2 m_W^4 \Gamma_\mu} \left| \sum_k K_{ek} K_{\mu k}^* G_\gamma \left(\frac{m_{\nu k}^2}{m_W^2} \right) \right|^2$$

↑
Rectangular matrix

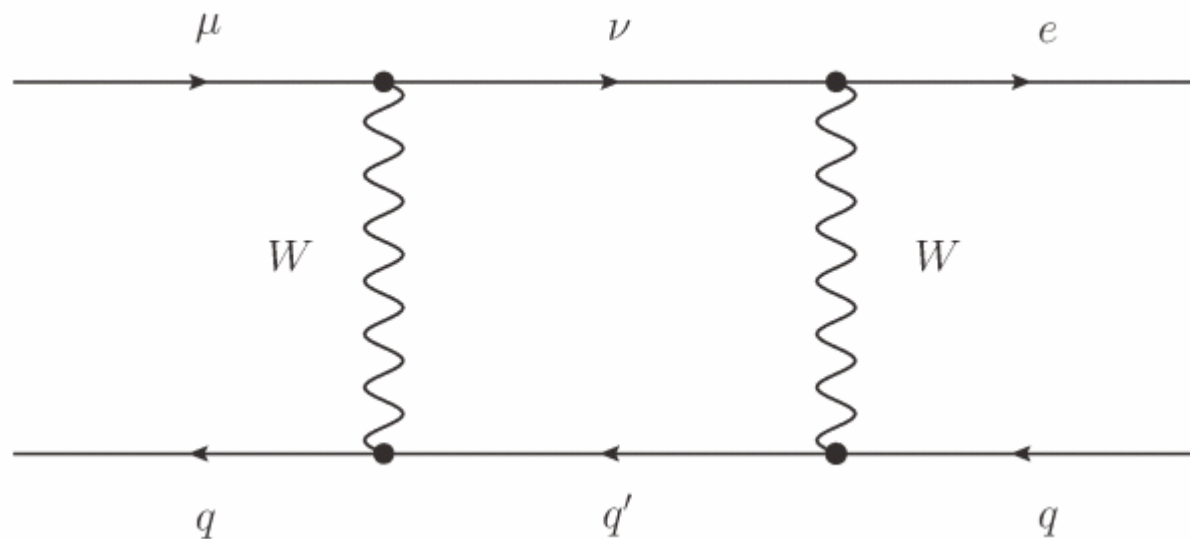


The **GIM**
suppression is
spoiled by the
singlet neutrinos

Much larger rates expected!

Boxes in the inverse seesaw

Furthermore, for $\mu - e$ conversion in nuclei and $l_i \rightarrow 3 l_j \dots$



[Ilakovac, Pilaftsis, 2009; Dinh, Ibarra, Molinaro, Petcov, 2012; Alonso, Dhen, Gavela, Hambye, 2013; Ilakovac, Pilaftsis, Popov, 2012]

- **Non-supersymmetric** contribution
- Relevant for **light singlet neutrinos**
- Large **non-dipole** contributions

Low-scale seesaw models

[Abada, Krauss, Porod, Staub, AV, Weiland, 2014]

75 pages paper

First complete study of all SUSY and non-SUSY contributions!

- Analytical and numerical study of

$$\ell_i \rightarrow \ell_j \gamma$$

$$\ell_i \rightarrow \ell_j \ell_k \ell_k$$

$\mu - e$ conversion in nuclei

- All contributions included
- A few hundred Feynman diagrams



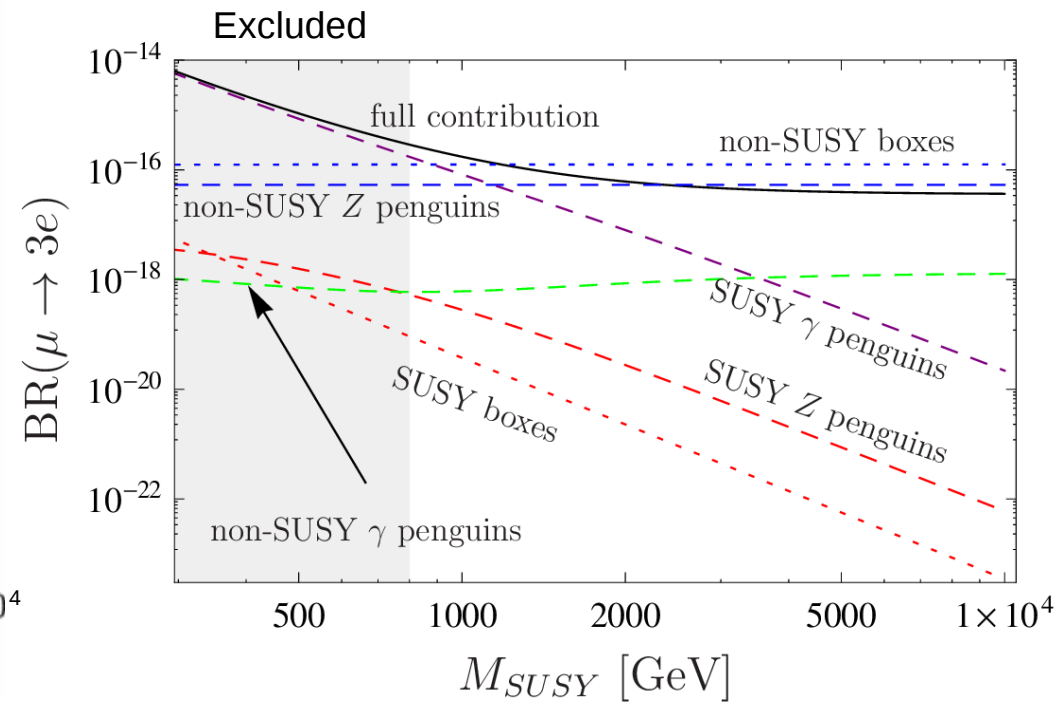
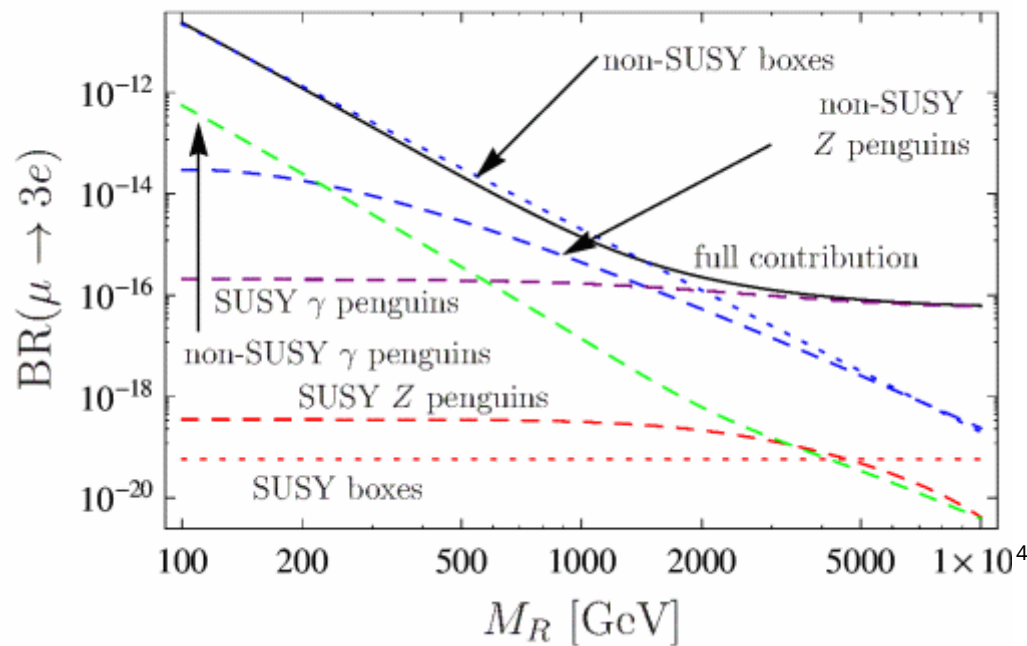
How were they
computed?

Later...

Low-scale seesaw models

$$l_i \rightarrow 3 l_j$$

[Abada, Krauss, Porod, Staub, AV, Weiland, 2014]



The anatomy of LFV strongly depends on M_R and M_{SUSY}

FlavorKit

[Porod, Staub, AV, 2014]

A computer tool that provides automatized analytical and numerical computation of flavor observables. It is based on **SARAH**, **SPheno** and **FeynArts/FormCalc**.

Lepton flavor	Quark flavor
$l_\alpha \rightarrow l_\beta \gamma$	$B_{s,d}^0 \rightarrow l^+ l^-$
$l_\alpha \rightarrow 3 l_\beta$	$\bar{B} \rightarrow X_s \gamma$
$\mu - e$ conversion in nuclei	$\bar{B} \rightarrow X_s l^+ l^-$
$\tau \rightarrow P l$	$\bar{B} \rightarrow X_{d,s} \nu \bar{\nu}$
$h \rightarrow l_\alpha l_\beta$	$B \rightarrow K l^+ l^-$
$Z \rightarrow l_\alpha l_\beta$	$K \rightarrow \pi \nu \bar{\nu}$
	$\Delta M_{B_{s,d}}$
	ΔM_K and ε_K
	$P \rightarrow l \nu$

Not limited to a single model: use it for the **model of your choice**

Easily **extendable**

Many observables ready to be computed in your favourite model!

Manual: [arXiv:1405.1434](https://arxiv.org/abs/1405.1434)

Website: <http://sarah.hepforge.org/FlavorKit.html>

LFV in B-meson decays

**With Sofiane M. Boucenna, Paulina Rocha-Morán
and José W. F. Valle**

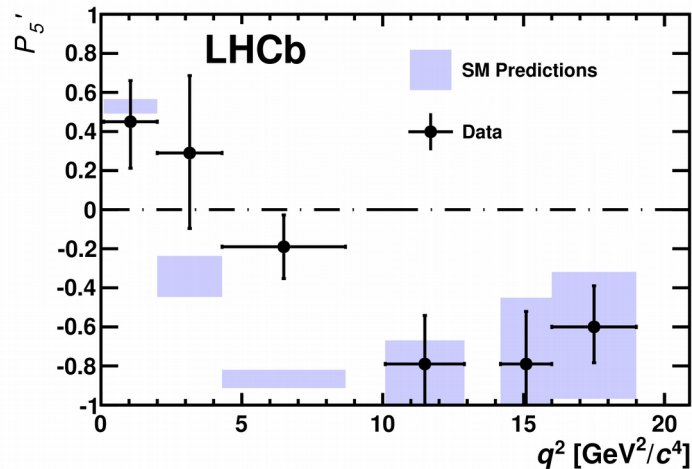
[PLB 750 \(2015\) 367 \[arXiv:1503.07099\]](#)

[arXiv:1810.02135](#)

The $b \rightarrow s$ anomalies

Episode IV: A new hope

2013 : First anomalies found by LHCb



Episode VI: Return of the anomalies

2015 : LHCb confirms first anomalies

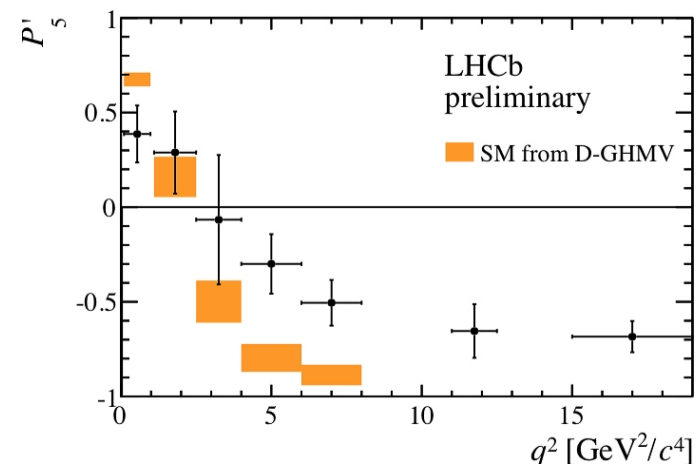
Episode V: LHCb strikes back

2014 : Lepton universality violation

$$R_K = \frac{\text{BR}(B \rightarrow K \mu^+ \mu^-)}{\text{BR}(B \rightarrow K e^+ e^-)} = 0.745_{-0.074}^{+0.090} \pm 0.036$$

$$R_K^{\text{SM}} \sim 1.00 \pm 0.01$$

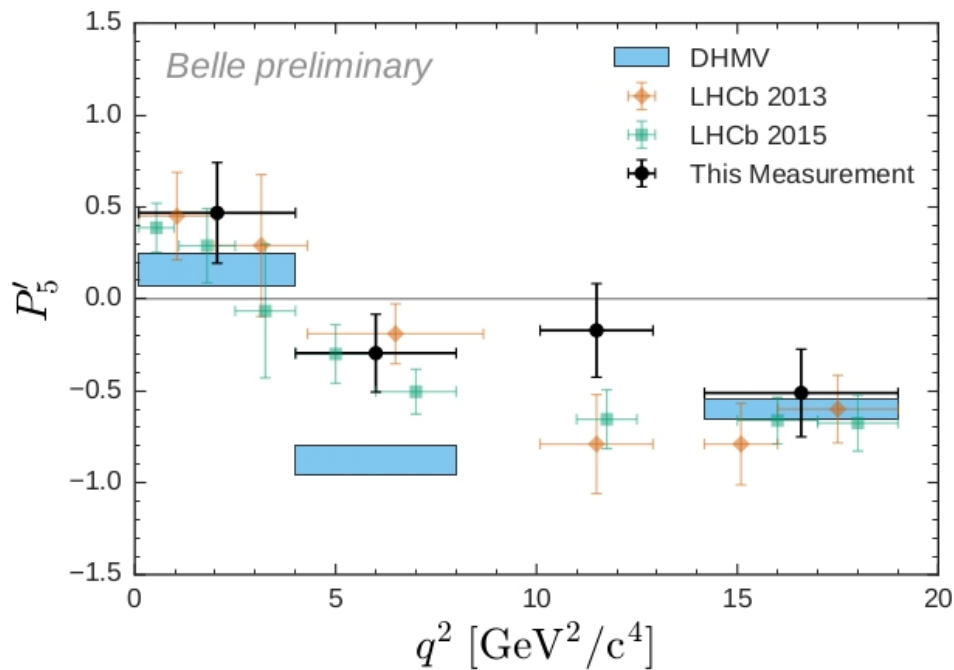
2.6σ away from the SM



The $b \rightarrow s$ anomalies

Episode I: The Belle menace

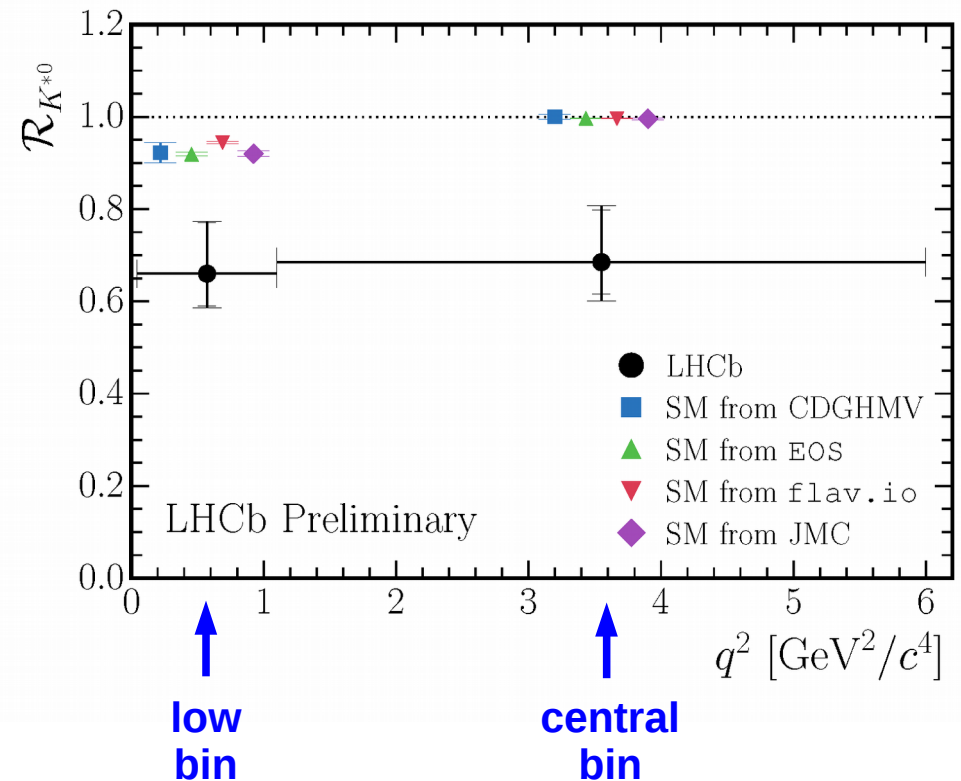
2016 : Belle finds additional hints



P'_5 anomaly confirmed
+ little LFVU indication

Episode II: Attack of R_{K^*}

2017 : More universality violation in LHCb



[No new episode in 2018 though....]

The $b \rightarrow s$ anomalies

Beyond the Standard Model



The $b \rightarrow s$ anomalies

Boring

Sizable corrections

(+ fluctuations)



LFV in B meson decays

What about LFV?

LFV in B meson decays

What about LFV?

[Glashow et al, 2014]

Lepton universality violation generically implies lepton flavor violation

Gauge basis

Mass basis

$$\mathcal{O} = \tilde{C}^Q (\bar{q}' \gamma_\alpha P_L q') \tilde{C}^L (\bar{\ell}' \gamma^\alpha P_L \ell') \longrightarrow \mathcal{O} = C^Q (\bar{q} \gamma_\alpha P_L q) C^L (\bar{\ell} \gamma^\alpha P_L \ell)$$

$$C^L = U_\ell^\dagger \tilde{C}^L U_\ell$$

However: we must have a **flavor theory** in order to make **predictions**

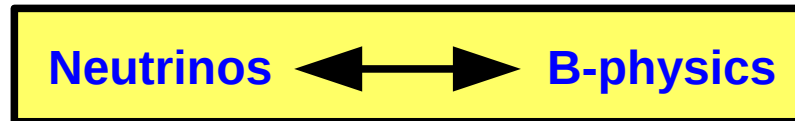
Are the LHCb anomalies related to neutrino oscillations?

Working hypothesis: What if $U_\ell = K^\dagger$?

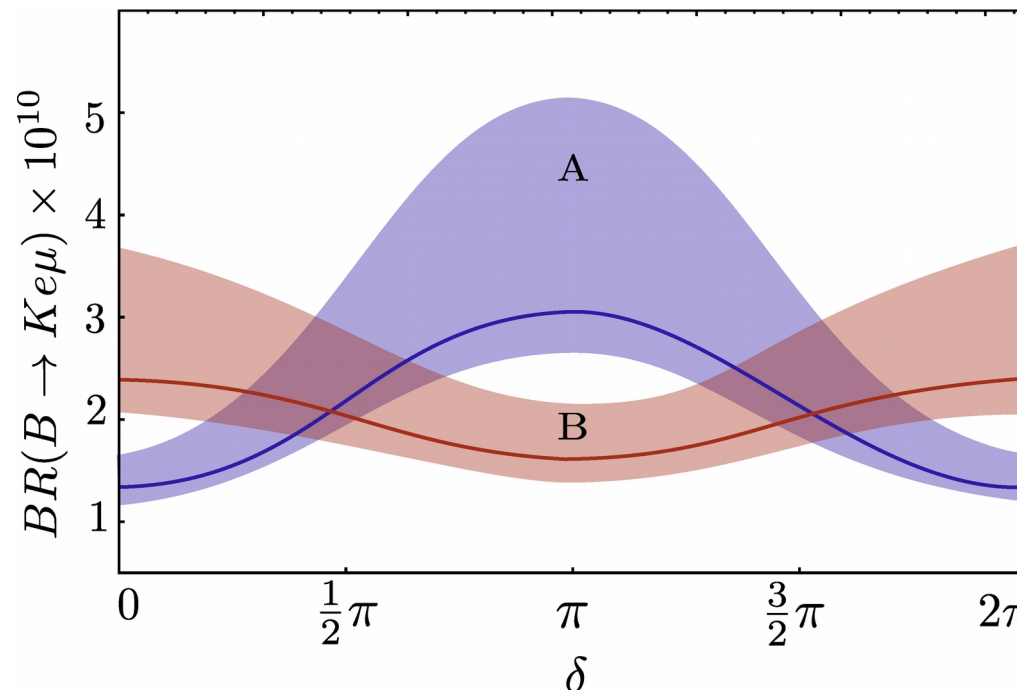
[Boucenna, Valle, AV, 2015]



Neutrino oscillations



Non-trivial
link!



Lines: BF
Bands: 1σ

An explicit model

[Rocha-Moran, AV, 2018]
 Extension of [Aristizabal Sierra, Staub, AV, 2015]

	generations	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$	$U(1)_X$	
H	1	1	2	1/2	0	} $U(1)_X$ breaking: $m_{Z'}$
ϕ	1	1	1	0	2	
S	1	1	1	0	-4	
q_L	3	3	2	1/6	0	} Vector-like
u_R	3	3	1	2/3	0	
d_R	3	3	1	-1/3	0	
ℓ_L	3	1	2	-1/2	0	
e_R	3	1	1	-1	0	
$Q_{L,R}$	1	3	2	1/6	2	
$L_{L,R}$	2	1	2	-1/2	2	
$F_{L,R}$	2	1	1	0	2	

Gauge symmetry
 Z' boson

$U(1)_X$ breaking: $m_{Z'}$

Vector-like

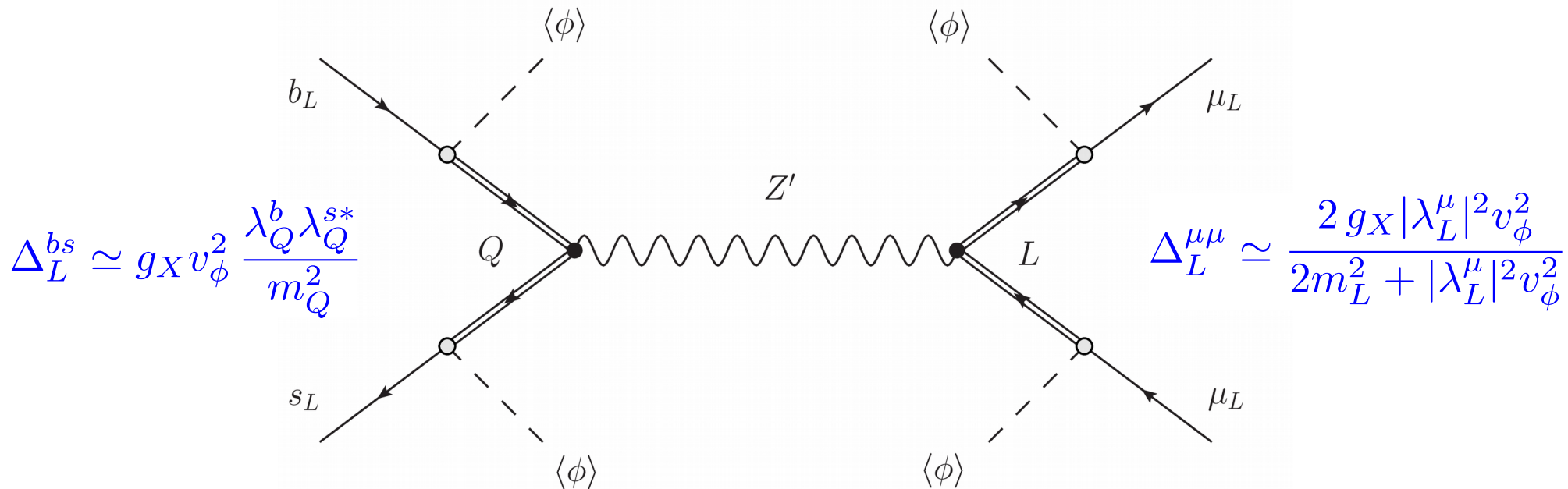
Vector-like = "joker"
 for model builders



Solving the $b \rightarrow s$ anomalies

[Aristizabal Sierra, Staub, AV, 2015]

Similar to
Altmannshofer et al,
Crivellin et al, 2014



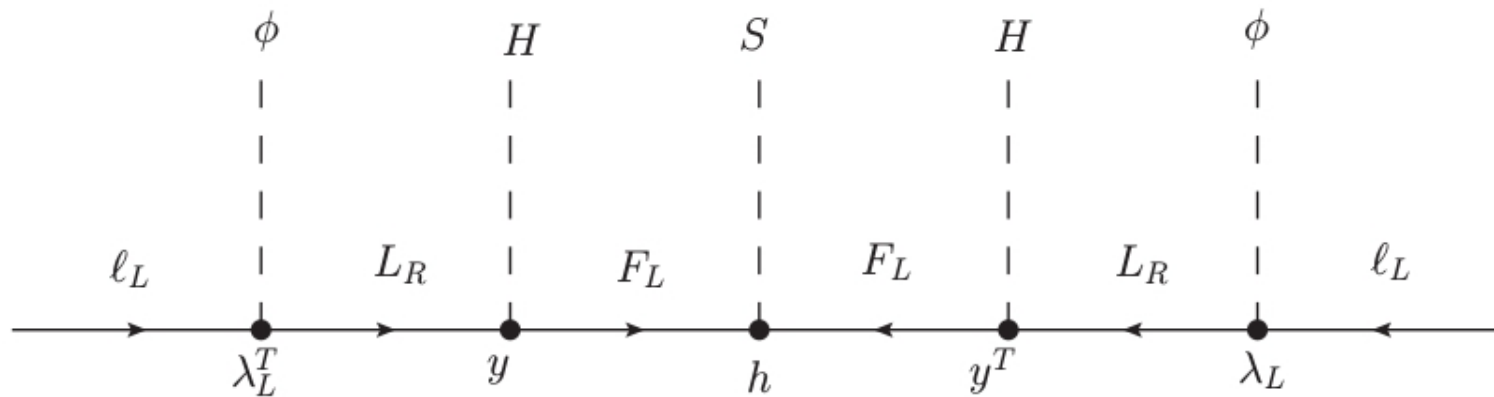
$$\mathcal{O} = (\bar{s} \gamma_\alpha P_L b) (\bar{\mu} \gamma^\alpha P_L \mu)$$

$$C_9^{\text{NP}} = -C_{10}^{\text{NP}}$$

Z' couplings induced by mixing with
vector-like fermions

Neutrino mass generation

[Rocha-Morán, AV, 2018]



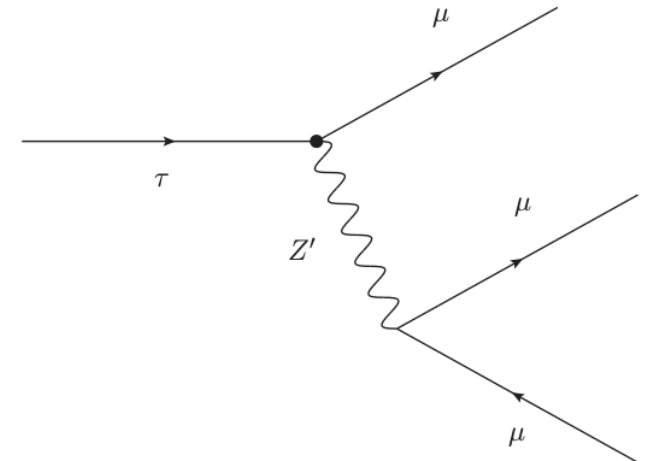
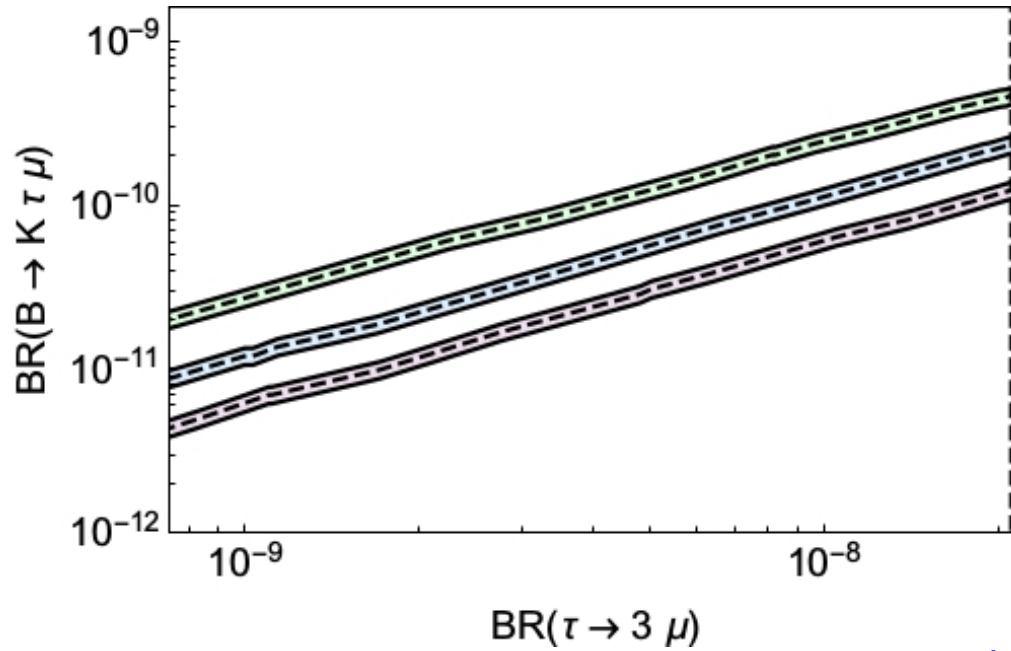
$$m \simeq \frac{v^2 v_\phi^2 v_S}{2\sqrt{2}} \lambda_L^T m_L^{-1} y m_F^{-1} h (m_F^{-1})^T y^T (m_L^{-1})^T \lambda_L$$

Inverse seesaw-like mechanism

$$h \ll 1$$

LFV phenomenology

[Rocha-Morán, AV, 2018]



Bs-mixing

$$\frac{\text{BR}(B \rightarrow K\tau\mu)}{\text{BR}(\tau \rightarrow 3\mu)} = 1.7 \cdot 10^7 \text{ TeV}^4 \left(\frac{|\Delta_L^{bs}|}{m_{Z'}} \right)^4 \frac{1}{|C_9^{\mu\mu, \text{NP}}|^2}$$

$$\text{BR}(B \rightarrow K\tau\mu)_{\text{max}} \lesssim 8 \cdot 10^{-10}$$

Anomaly

Correlation is (almost)* unavoidable

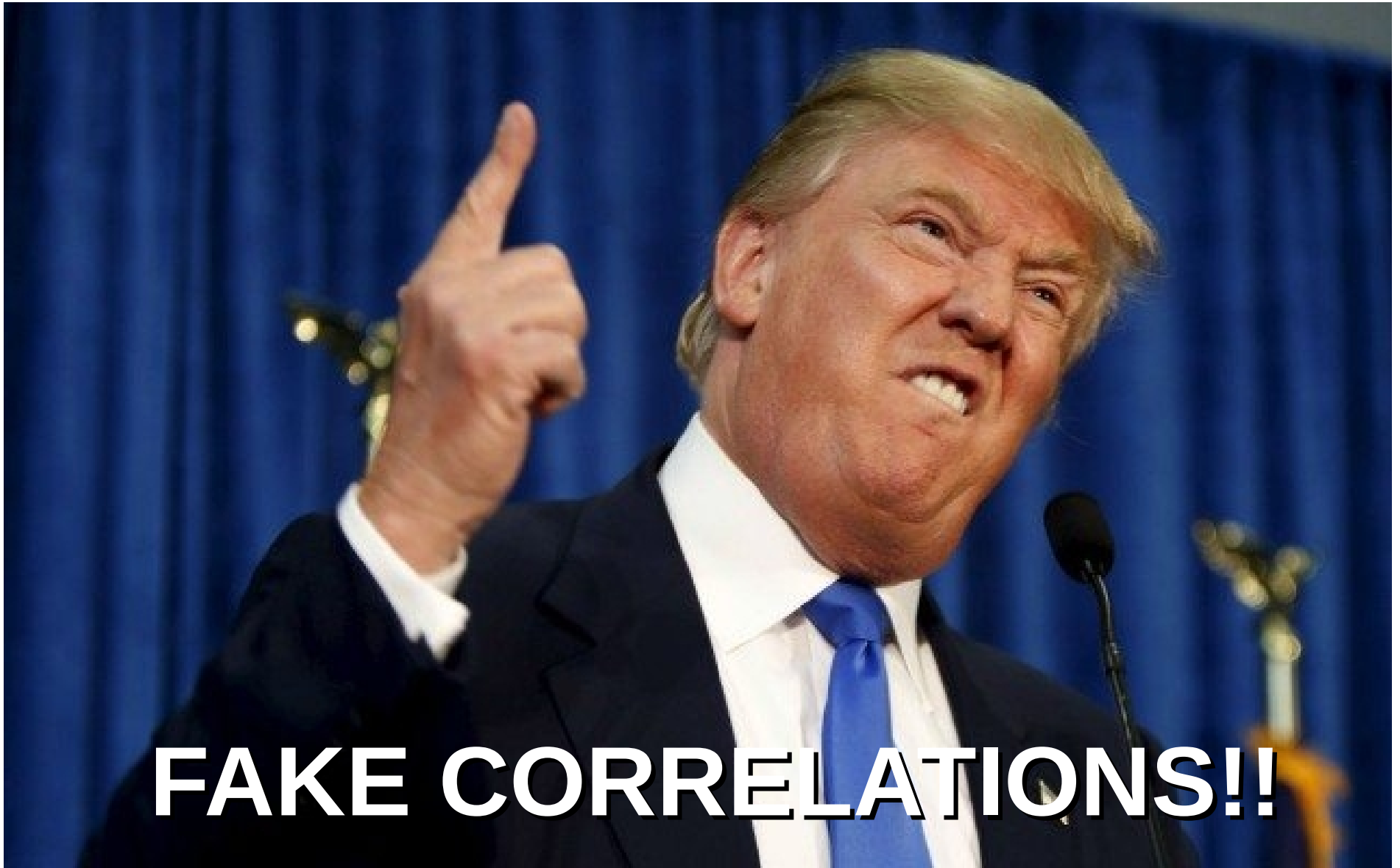
* : could be broken by loops (indeed possible!)

Master Majorana parametrization

With Isabel Cordero-Carrión and Martin Hirsch

[arXiv:1812.03896](https://arxiv.org/abs/1812.03896)

Motivation



The master formula

All Majorana
neutrino mass
models!

$$m = f (y_1^T M y_2 + y_2^T M^T y_1)$$

m : symmetric 3×3 Majorana neutrino mass matrix

f : global factor [numerical factors, model parameters, mass ratios, ...]

y_1 : $n_1 \times 3$ Yukawa matrix

y_2 : $n_2 \times 3$ Yukawa matrix [$n_1 \geq n_2$]

M : $n_1 \times n_2$ matrix [dimensions of mass]

Towards a master parametrization

$$m = f \left(y_1^T M y_2 + y_2^T M^T y_1 \right)$$

↑
Neutrino oscillation experiments

↑
Model parameters

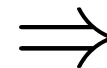
Goal:

To establish a general, complete and programmable parametrization of the y_1 and y_2 Yukawa matrices

Particular case: Type-I seesaw

Casas-Ibarra parametrization

[Casas, Ibarra, 2001]



Master parametrization

The master parametrization

$$y_1 = \frac{1}{\sqrt{2}f} V_1^\dagger \begin{pmatrix} \Sigma^{-1/2} W A \\ X_1 \\ X_2 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$y_2 = \frac{1}{\sqrt{2}f} V_2^\dagger \begin{pmatrix} \Sigma^{-1/2} \widehat{W}^* \widehat{B} \\ X_3 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

Summary:

	Experimental input:	Model input:	Free Yukawa parameters:	
v-data	$\left\{ \begin{array}{l} \bar{D}_{\sqrt{m}} \\ U \end{array} \right\}$	$M \left\{ \begin{array}{l} f \\ \Sigma \\ V_{1,2} \end{array} \right\}$	$\left\{ \begin{array}{l} \widehat{W} \\ X_{1,2,3} \\ T \\ K \\ (C_2) \end{array} \right\}$	$\left\{ \begin{array}{l} A \\ \widehat{B} \end{array} \right\}$

The master parametrization

Matrix	Dimensions	Property	Real parameters
X_1	$(n_2 - n) \times 3$	Absent if $n = n_2$	$6(n_2 - n)$
X_2	$(n_1 - n_2) \times 3$	Absent if $n_1 = n_2$	$6(n_1 - n_2)$
X_3	$(n_2 - n) \times 3$	Absent if $n = n_2$	$6(n_2 - n)$
W	$n \times r$		$r(2n - r)$
T	$r \times r$	Upper triangular with $(T)_{ii} > 0$	r^2
K	$r \times r$	Antisymmetric	$r(r - 1)$
\bar{B}	$(n - r) \times 3$	Absent if $n = r$	$6(n - r)$
C_1	$r \times 3$	Case-dependent	0 or 2
C_2	3×3	Case-dependent	-

$$\#_{\text{free}} = \#_{X_1} + \#_{X_2} + \#_{X_3} + \#_T + \#_W + \#_K + \#\bar{B} + \#_{C_1}$$

BNT model

	generations	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$
Φ	1	1	4	3/2
$\psi_{L,R}$	3	1	3	-1

$$\Phi = \begin{pmatrix} \Phi^{+++} \\ \Phi^{++} \\ \Phi^+ \\ \Phi^0 \end{pmatrix}$$

$$-\mathcal{L} \supset y_\psi \bar{L} H \psi_R + y_{\bar{\psi}} \bar{L}^c \Phi \psi_L + M_\psi \bar{\psi} \psi$$

An example of $y_1 \neq y_2$

$$\psi_{L,R} = \begin{pmatrix} \psi^0 \\ \psi^- \\ \psi^{--} \end{pmatrix}_{L,R}$$

$$\mathcal{V} \supset \lambda_\Phi H^3 \Phi$$

\Rightarrow

$$\langle \Phi^0 \rangle = \frac{\lambda_\Phi v^3}{2\sqrt{2}M_\Phi^2}$$

\Rightarrow

$$\Delta L = 2$$

Lepton number violation

[Babu, Nandi, Tavartkiladze, 2009]

BNT model

$$m = f (y_1^T M y_2 + y_2^T M^T y_1)$$

$$m = \frac{\lambda_\Phi v^4}{4 M_\Phi^2} \left[y_\psi^T M_\psi^{-1} y_{\bar{\psi}} + y_{\bar{\psi}}^T (M_\psi^{-1})^T y_\psi \right]$$

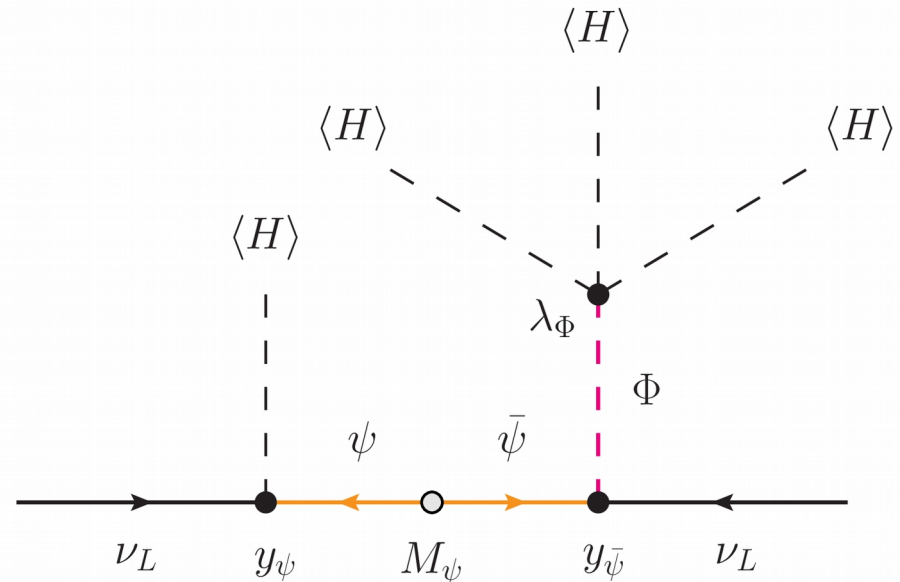
$$f = \frac{\lambda_\Phi v^2}{2 M_\Phi^2}$$

$$n_1 = n_2 = 3$$

$$y_1 = y_\psi \neq y_2 = y_{\bar{\psi}}$$

$$M = \frac{v^2}{2} M_\psi^{-1}$$

\Rightarrow

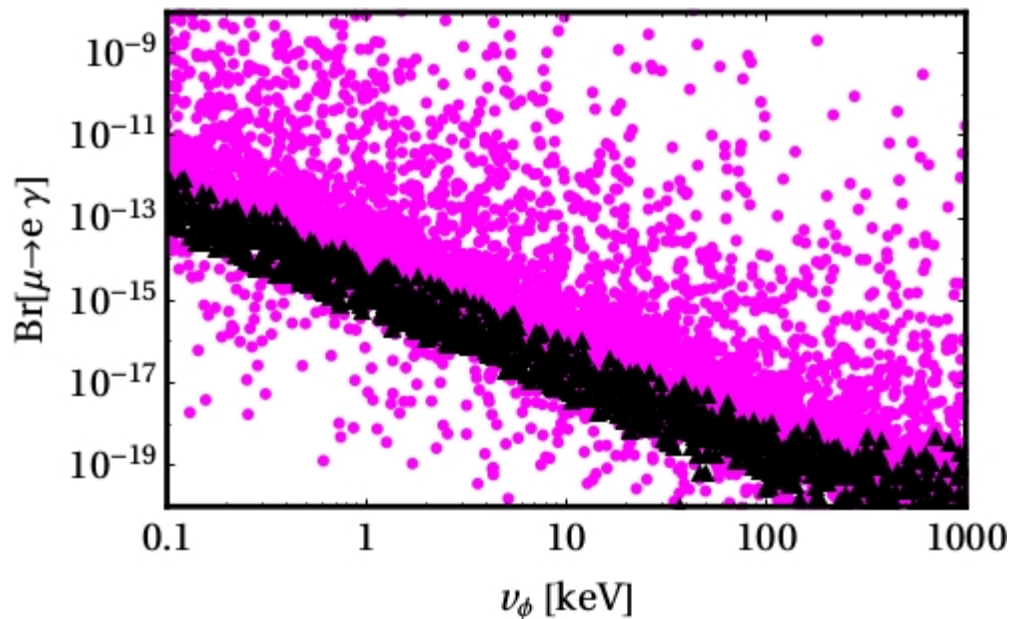


[Babu, Nandi, Tavartkiladze, 2009]

ν -data NH within 3σ

$$M_\psi \in [0.5, 2] \text{ TeV}$$

$$W = \mathbb{I}$$



Black: Trivial scan [$T = \mathbb{I}$ & $K = 0$]

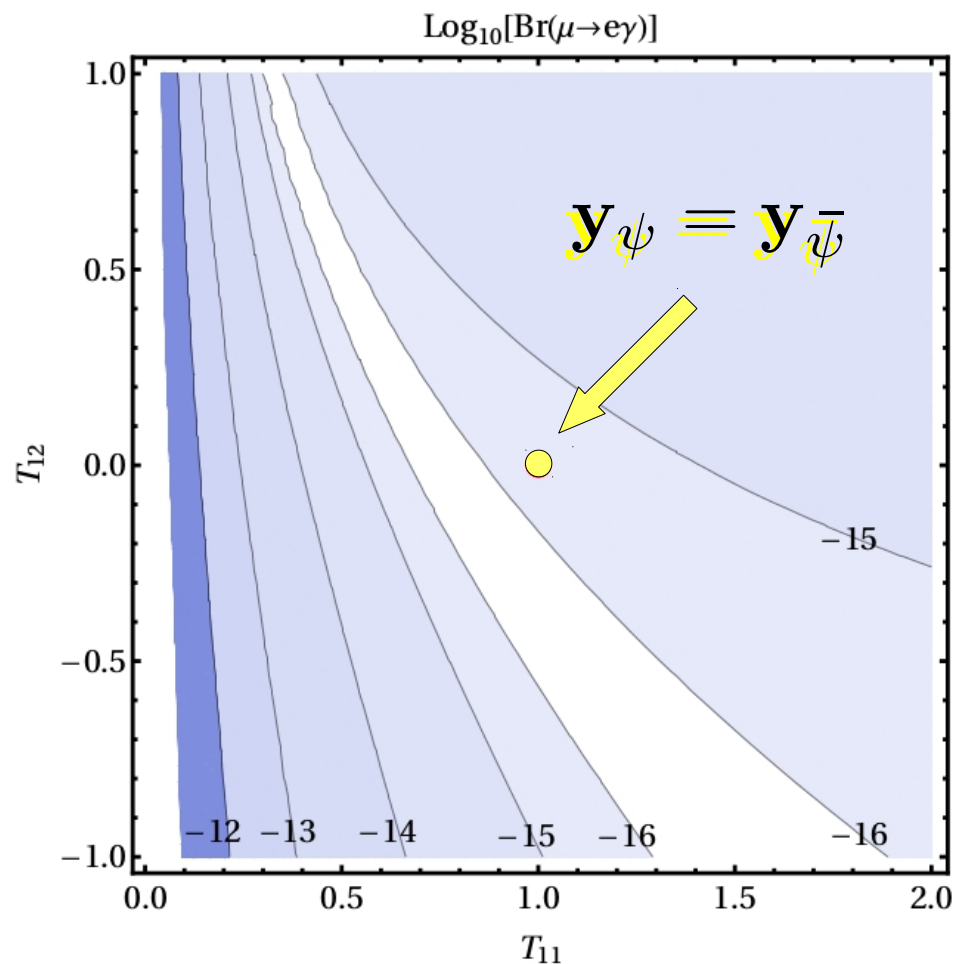
Purple: General scan

A large parameter space that can only be covered with the **master parametrization**

ν -data NH BFP

$$v_\Phi = 10^{-5} \text{ GeV}$$

$$M_\psi = 0.5 \text{ TeV}$$



Model in **SARAH** and scans with **SSP** [Staub]
BR computed with **FlavorKit** [Porod, Staub, AV]

Final remarks

Final remarks

LFV is going to live a **golden age**

Many LFV observables. **Correlations** are not only possible, but in fact expected!

We must be **ready**: understand the LFV anatomy, patterns, correlations, hierarchies...

Thank you!

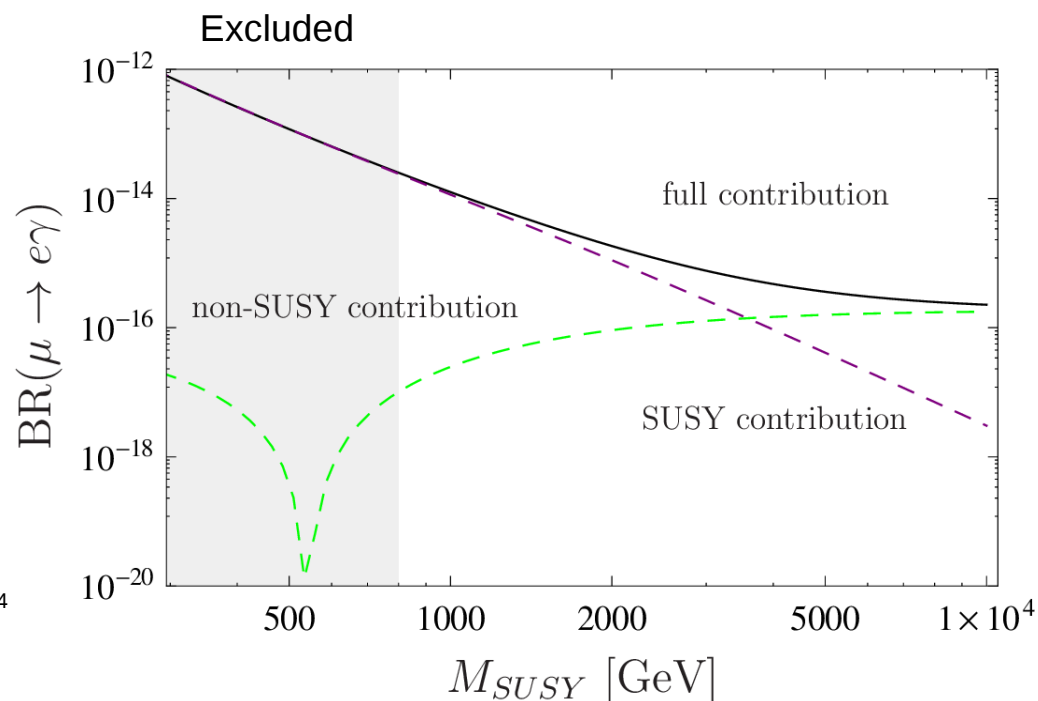
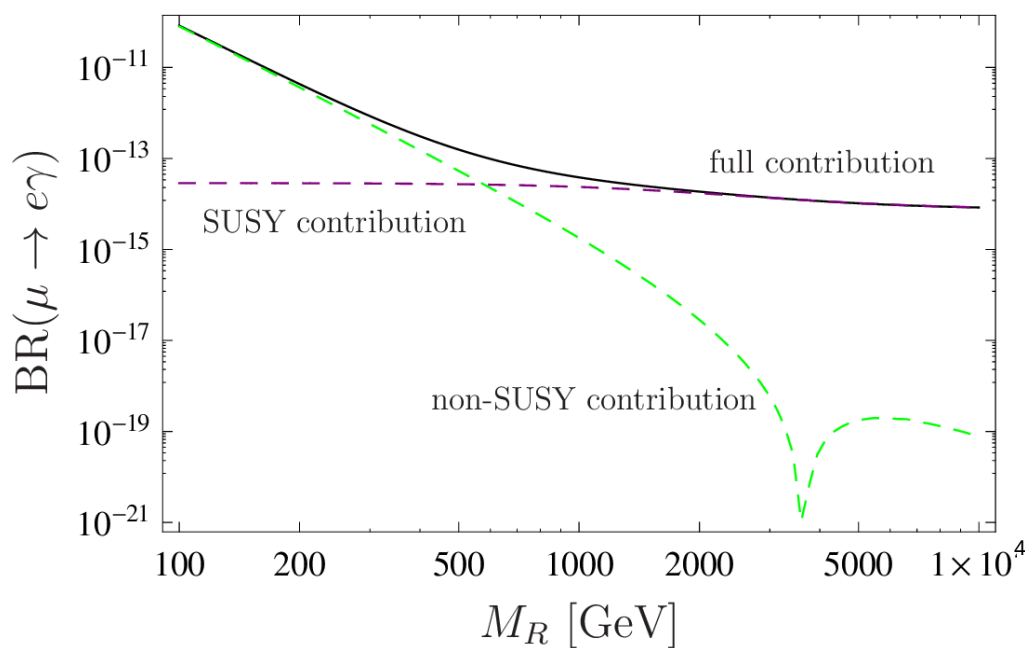


Backup slides

Low-scale seesaw models

$$l_i \rightarrow l_j \gamma$$

[Abada, Krauss, Porod, Staub, AV, Weiland, 2014]

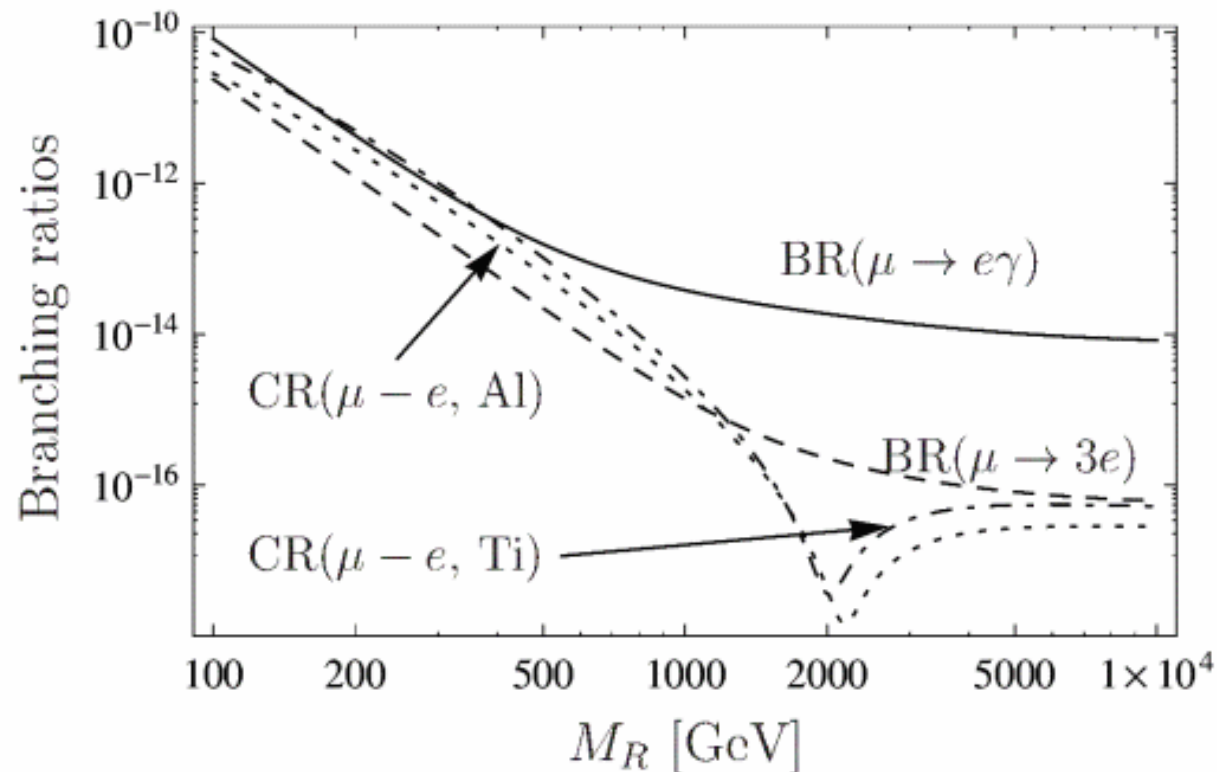


The anatomy of LFV strongly depends on M_R and M_{SUSY}

Low-scale seesaw models

$$l_i \rightarrow 3 l_j$$

[Abada, Krauss, Porod, Staub, AV, Weiland, 2014]

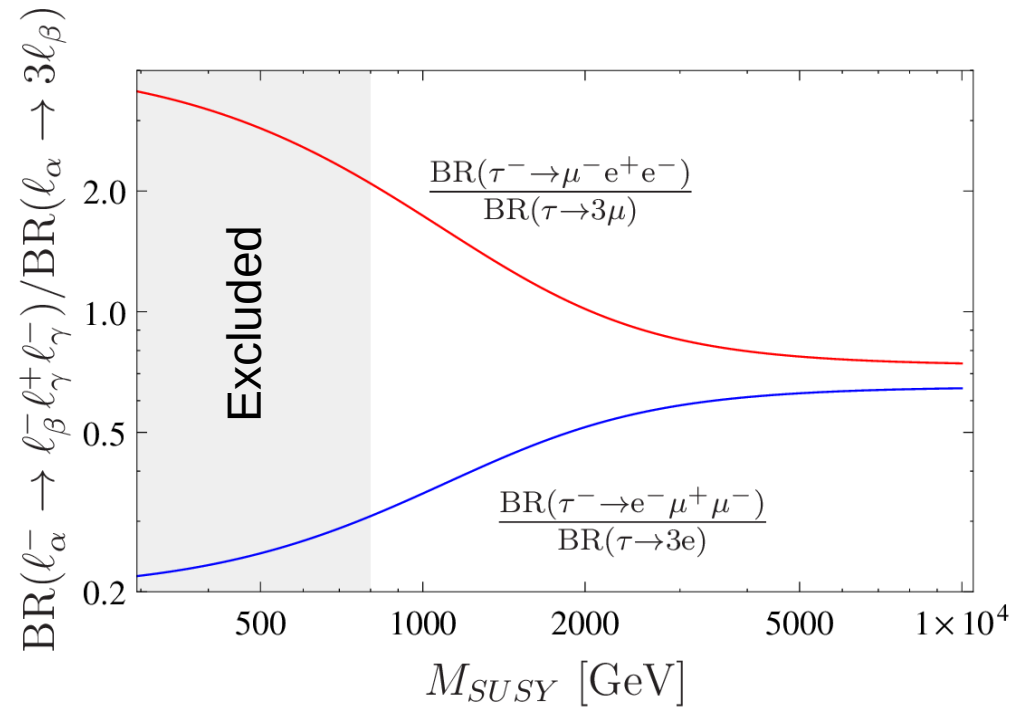
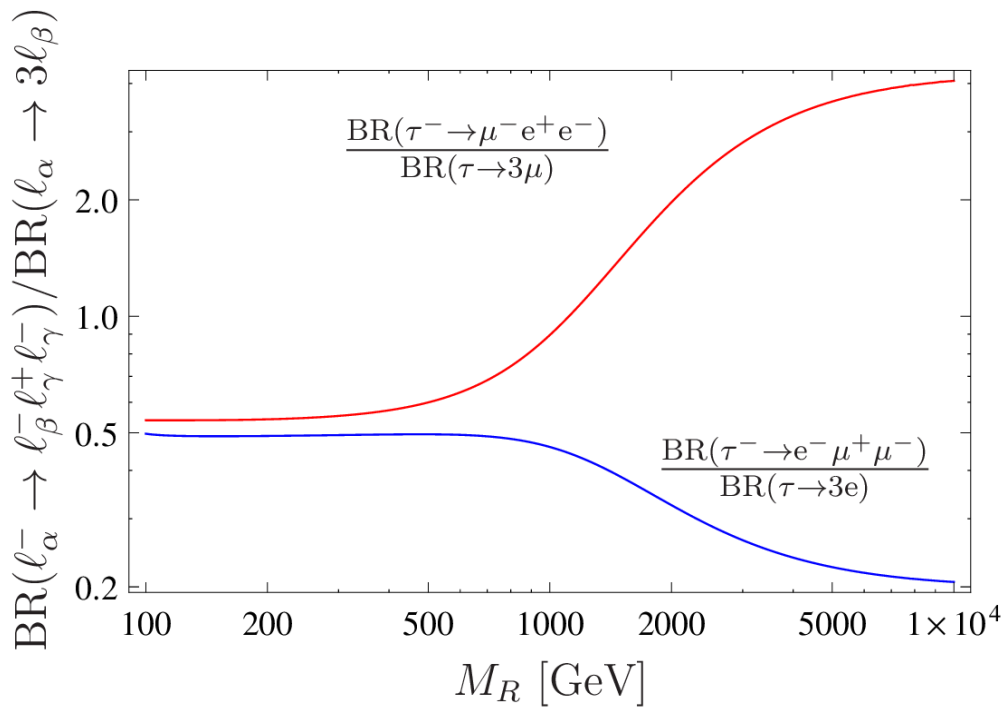


The **dipole dominance** is broken for low RH neutrino masses

Low-scale seesaw models

$$l_i \rightarrow l_j l_k l_k$$

[Abada, Krauss, Porod, Staub, AV, Weiland, 2014]



Tau LFV decay ratios provide information on the **mass scales**

Interpreting the anomalies

$$\boxed{b \rightarrow s}$$

Effective hamiltonian

$$\mathcal{H}_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} V_{tb}V_{ts}^* \frac{e^2}{16\pi^2} \sum_i (C_i \mathcal{O}_i + C'_i \mathcal{O}'_i) + \text{h.c.}$$

C_i : Wilson coefficients

\mathcal{O}_i : Operators

$$\mathcal{O}_9 = (\bar{s}\gamma_\mu P_L b) (\bar{\ell}\gamma^\mu \ell)$$

$$\mathcal{O}'_9 = (\bar{s}\gamma_\mu P_R b) (\bar{\ell}\gamma^\mu \ell)$$

$$\mathcal{O}_{10} = (\bar{s}\gamma_\mu P_L b) (\bar{\ell}\gamma^\mu \gamma_5 \ell)$$

$$\mathcal{O}'_{10} = (\bar{s}\gamma_\mu P_R b) (\bar{\ell}\gamma^\mu \gamma_5 \ell)$$

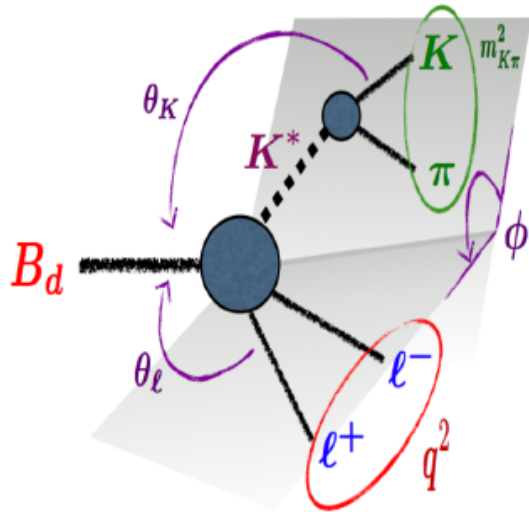
$$C_i = C_i^{\text{SM}} + C_i^{\text{NP}}$$

[analogous for primed operators]

The $b \rightarrow s$ anomalies

$B \rightarrow K^* (\rightarrow K \pi) \mu^+ \mu^-$ differential angular distribution

$$\frac{d^4\Gamma}{dq^2 d\cos\theta_K d\cos\theta_l d\phi} = \frac{9}{32\pi} \left[J_{1s} \sin^2\theta_K + J_{1c} \cos^2\theta_K + (J_{2s} \sin^2\theta_K + J_{2c} \cos^2\theta_K) \cos 2\theta_l \right. \\ \left. + J_3 \sin^2\theta_K \sin^2\theta_l \cos 2\phi + J_4 \sin 2\theta_K \sin 2\theta_l \cos\phi + J_5 \sin 2\theta_K \sin\theta_l \cos\phi \right. \\ \left. + (J_{6s} \sin^2\theta_K + J_{6c} \cos^2\theta_K) \cos\theta_l + J_7 \sin 2\theta_K \sin\theta_l \sin\phi \right. \\ \left. + J_8 \sin 2\theta_K \sin 2\theta_l \sin\phi + J_9 \sin^2\theta_K \sin^2\theta_l \sin 2\phi \right]$$



[Figure borrowed from Javier Virto]

J_i : functions of q^2 , C_i , FF

Optimized observables
[Descotes-Genon et al, 2012, 2013]

$$P'_5 = \frac{J_5}{2\sqrt{-J_{2s}J_{2c}}}$$

LFV at LHCb

Lepton flavor violating decays at



$$B_{d,s}^0 \rightarrow \ell_i \ell_j$$

[Aaij et al, LHCb collaboration, 2013]

$$\tau \rightarrow 3\mu$$

[Aaij et al, LHCb collaboration, 2014]

Impressive result in a
hadronic machine

Limits **improved** with respect to **CDF**

$$\text{BR}(B^0 \rightarrow e\mu) < 2.8 \cdot 10^{-9}$$

$$\text{BR}(B_s^0 \rightarrow e\mu) < 1.1 \cdot 10^{-8}$$

Large production of τ 's, clean **final state**

$$\text{BR}(\tau \rightarrow 3\mu) < 4.6 \cdot 10^{-8} \text{ (at 90\% CL)}$$

To be compared with $2.1 \cdot 10^{-8}$ (**Belle**)

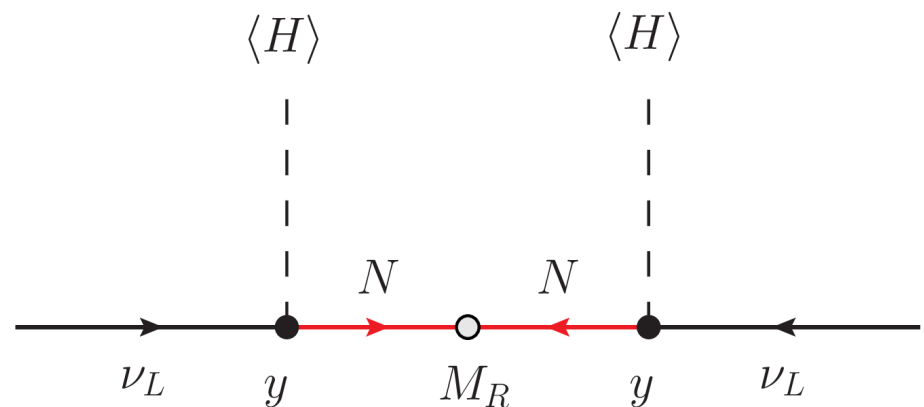
Type-I seesaw

everyone's
model

$$m = f (y_1^T M y_2 + y_2^T M^T y_1)$$

$$\left. \begin{aligned} f &= -1 \\ n_1 &= n_2 = 3 \\ y_1 &= y_2 = y/\sqrt{2} \\ M &= \frac{v^2}{2} M_R^{-1} \end{aligned} \right\} \Rightarrow$$

$$m = -\frac{v^2}{2} y^T M_R^{-1} y$$



[Minkowski, 1977]

Inverse seesaw

$$m = f (y_1^T M y_2 + y_2^T M^T y_1)$$

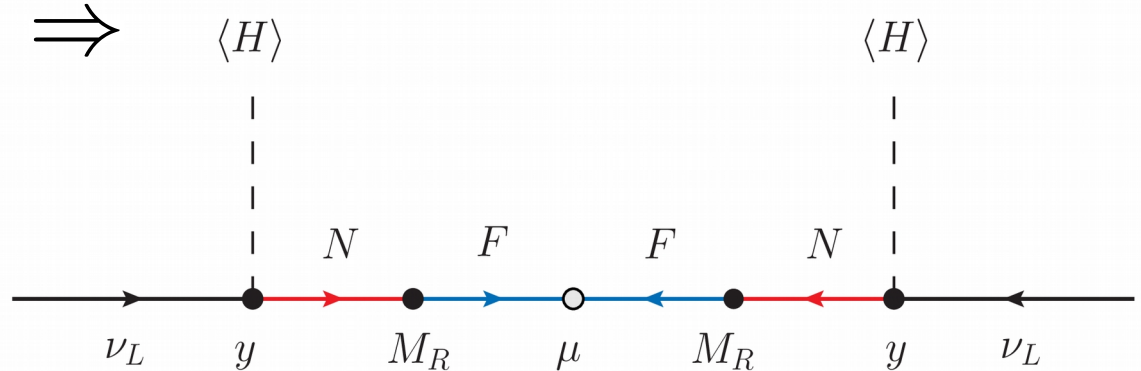
$$f = 1$$

$$n_1 = n_2 = 3$$

$$y_1 = y_2 = y/\sqrt{2}$$

$$M = \frac{v^2}{2} (M_R^T)^{-1} \mu M_R^{-1}$$

$$m = \frac{v^2}{2} y^T (M_R^T)^{-1} \mu M_R^{-1} y$$



[Mohapatra, Valle, 1986]

Scotogenic model

$$m = f (y_1^T M y_2 + y_2^T M^T y_1)$$

$$m = \frac{\lambda_5 v^2}{32\pi^2} y^T M_R^{-1} f_{\text{loop}} y$$

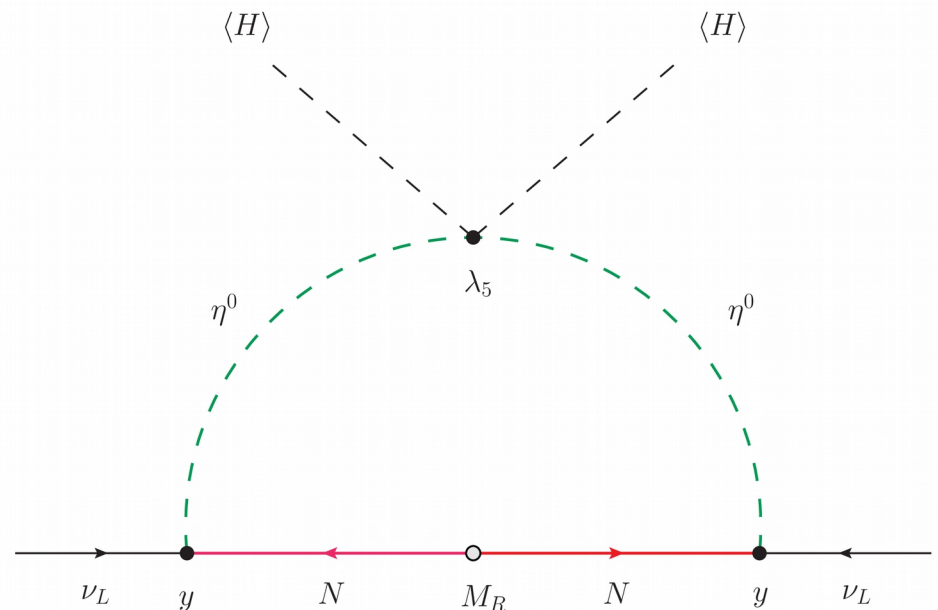
$$f = \frac{\lambda_5}{16\pi^2}$$

$$n_1 = n_2 = 3$$

$$y_1 = y_2 = y/\sqrt{2}$$

$$M = \frac{v^2}{2} M_R^{-1} f_{\text{loop}}$$

\Rightarrow



[Ma, 2006]

The master parametrization

$$y_1 = \frac{1}{\sqrt{2}f} V_1^\dagger \begin{pmatrix} \Sigma^{-1/2} W A \\ X_1 \\ X_2 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$y_2 = \frac{1}{\sqrt{2}f} V_2^\dagger \begin{pmatrix} \Sigma^{-1/2} \widehat{W}^* \widehat{B} \\ X_3 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$D_m = \text{diag}(m_1, m_2, m_3) = U^T m U$$

$$r_m = \text{rank}(m) \begin{cases} \bar{D}_{\sqrt{m}} = \text{diag}(\sqrt{m_1}, \sqrt{m_2}, \sqrt{m_3}) \\ \bar{D}_{\sqrt{m}} = \text{diag}(\sqrt{m_1}, \sqrt{m_2}, \sqrt{v}) \end{cases}$$

2 or 3

U : 3×3 unitary matrix

$$U^\dagger U = U U^\dagger = \mathbb{I}_3$$

Leptonic mixing matrix

The master parametrization

$$y_1 = \frac{1}{\sqrt{2f}} V_1^\dagger \begin{pmatrix} \Sigma^{-1/2} W A \\ X_1 \\ X_2 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$y_2 = \frac{1}{\sqrt{2f}} V_2^\dagger \begin{pmatrix} \Sigma^{-1/2} \widehat{W}^* \widehat{B} \\ X_3 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$M = V_1^T \widehat{\Sigma} V_2$$

Singular-value decomposition

$$\widehat{\Sigma} = \begin{pmatrix} \Sigma & 0 \\ 0 & 0_{n_2-n} \\ 0_{n_1-n_2} \end{pmatrix}$$

$$\widehat{\Sigma} : n_1 \times n_2$$

$$V_1 : n_1 \times n_1$$

$$V_2 : n_2 \times n_2$$

Unitary matrices

$$\Sigma : n \times n$$

$$\Sigma = \text{diag}(\sigma_1, \sigma_2, \dots, \sigma_n) \quad (\sigma_i > 0)$$

The master parametrization

$$y_1 = \frac{1}{\sqrt{2f}} V_1^\dagger \begin{pmatrix} \Sigma^{-1/2} W A \\ X_1 \\ X_2 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$y_2 = \frac{1}{\sqrt{2f}} V_2^\dagger \begin{pmatrix} \Sigma^{-1/2} \widehat{W}^* \widehat{B} \\ X_3 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$X_1 : (n_2 - n) \times 3$$

$$X_2 : (n_1 - n_2) \times 3$$

$$X_3 : (n_2 - n) \times 3$$

Arbitrary matrices

[Note: absent if $n_1 = n_2 = n$]

The master parametrization

$$y_1 = \frac{1}{\sqrt{2f}} V_1^\dagger \begin{pmatrix} \Sigma^{-1/2} W A \\ X_1 \\ X_2 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$y_2 = \frac{1}{\sqrt{2f}} V_2^\dagger \begin{pmatrix} \Sigma^{-1/2} \widehat{W}^* \widehat{B} \\ X_3 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$\widehat{W} = (W \quad \bar{W})$$

$$r = \text{rank}(W) \\ \leq \min(n, 3)$$

\widehat{W} : $n \times n$ unitary matrix

$$W : n \times r$$

$$\bar{W} : n \times (n - r) \longrightarrow \text{Absent if } n = r$$

$$\widehat{W}^\dagger \widehat{W} = \widehat{W} \widehat{W}^\dagger = \mathbb{I}_n$$

The master parametrization

$$y_1 = \frac{1}{\sqrt{2f}} V_1^\dagger \begin{pmatrix} \Sigma^{-1/2} W A \\ X_1 \\ X_2 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$
$$y_2 = \frac{1}{\sqrt{2f}} V_2^\dagger \begin{pmatrix} \Sigma^{-1/2} \widehat{W}^* \widehat{B} \\ X_3 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$A : r \times 3$$

$$A = T C_1$$

$$T : r \times r$$

Invertible upper triangular matrix

$$(T)_{ii} \in \mathbb{R}, (T)_{ii} > 0$$

$$C_1 : r \times 3$$

Numerical matrix whose form depends on r_m and r

The master parametrization

$$y_1 = \frac{1}{\sqrt{2f}} V_1^\dagger \begin{pmatrix} \Sigma^{-1/2} W A \\ X_1 \\ X_2 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$y_2 = \frac{1}{\sqrt{2f}} V_2^\dagger \begin{pmatrix} \Sigma^{-1/2} \widehat{W}^* \widehat{B} \\ X_3 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

$$\widehat{B} = \begin{pmatrix} B \\ \bar{B} \end{pmatrix}$$

$$B : r \times 3$$

$$B = (T^T)^{-1} [C_1 C_2 + K C_1]$$

$$C_2 : 3 \times 3$$

Matrix whose form depends on r_m and r

$$\widehat{B} : n \times 3$$

$$\bar{B} : (n - r) \times 3 \longrightarrow \begin{matrix} \text{Absent if} \\ n = r \end{matrix}$$

$K : r \times r$
antisymmetric

The master parametrization

$$y_1 = \frac{1}{\sqrt{2f}} V_1^\dagger \begin{pmatrix} \Sigma^{-1/2} W \mathbf{A} \\ X_1 \\ X_2 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$
$$y_2 = \frac{1}{\sqrt{2f}} V_2^\dagger \begin{pmatrix} \Sigma^{-1/2} \widehat{W}^* \widehat{\mathbf{B}} \\ X_3 \end{pmatrix} \bar{D}_{\sqrt{m}} U^\dagger$$

For $r_m = r = 3$:

$$C_1 = \mathbb{I}_3 \quad C_2 = \mathbb{I}_3 + K_{12} \frac{T_{13}}{T_{11}} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & -1 & 0 \end{pmatrix}$$

[Different matrix forms for other values of (r_m, r) . Ask me if you want to see them]

The Casas-Ibarra limit

Particular case: Type-I seesaw

$$m = -\frac{v^2}{2} y^T M_R^{-1} y$$

⇓

$$\left. \begin{aligned} f &= -1 \\ n_1 = n_2 = n &= 3 \\ r_m = r &= 3 \\ y_1 = y_2 &= y/\sqrt{2} \\ M &= \frac{v^2}{2} M_R^{-1} \end{aligned} \right\}$$

⇒

$$\left\{ \begin{aligned} V_1 = V_2 = V & \text{ (= } \mathbb{I} \text{ in mass basis)} \\ X_{1,2,3}, \bar{W} \text{ and } \bar{B} & \text{ are absent} \\ W^T W A = B & \Rightarrow B = (A^T)^{-1} \\ & \Rightarrow R = W A \\ & \text{orthogonal } 3 \times 3 \\ R^T R = R R^T &= \mathbb{I} \end{aligned} \right.$$

Casas-Ibarra parametrization [Casas, Ibarra, 2001]

$$y = i \Sigma^{-1/2} R D_{\sqrt{m}} U^\dagger$$

⇑